

ISMDB

InterStellar Medium DataBase

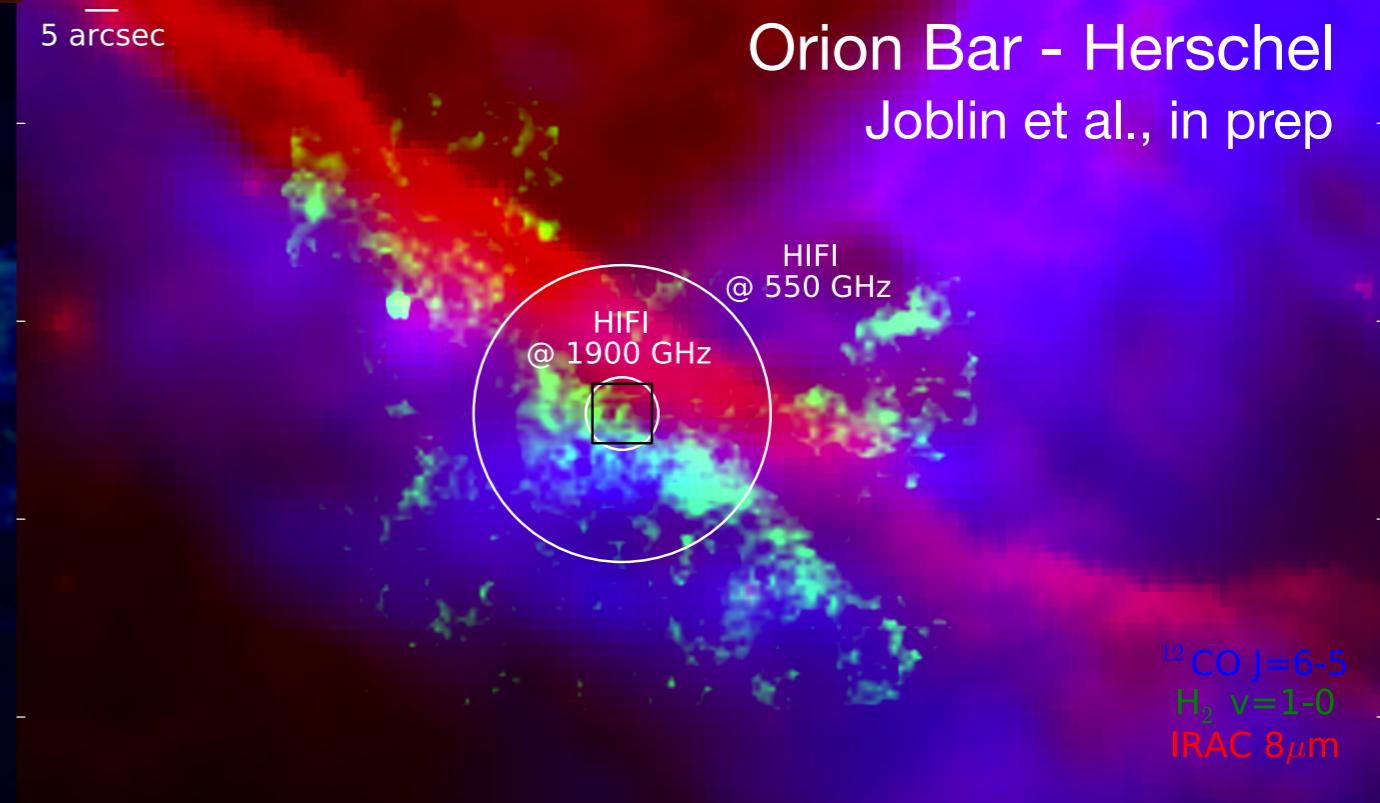
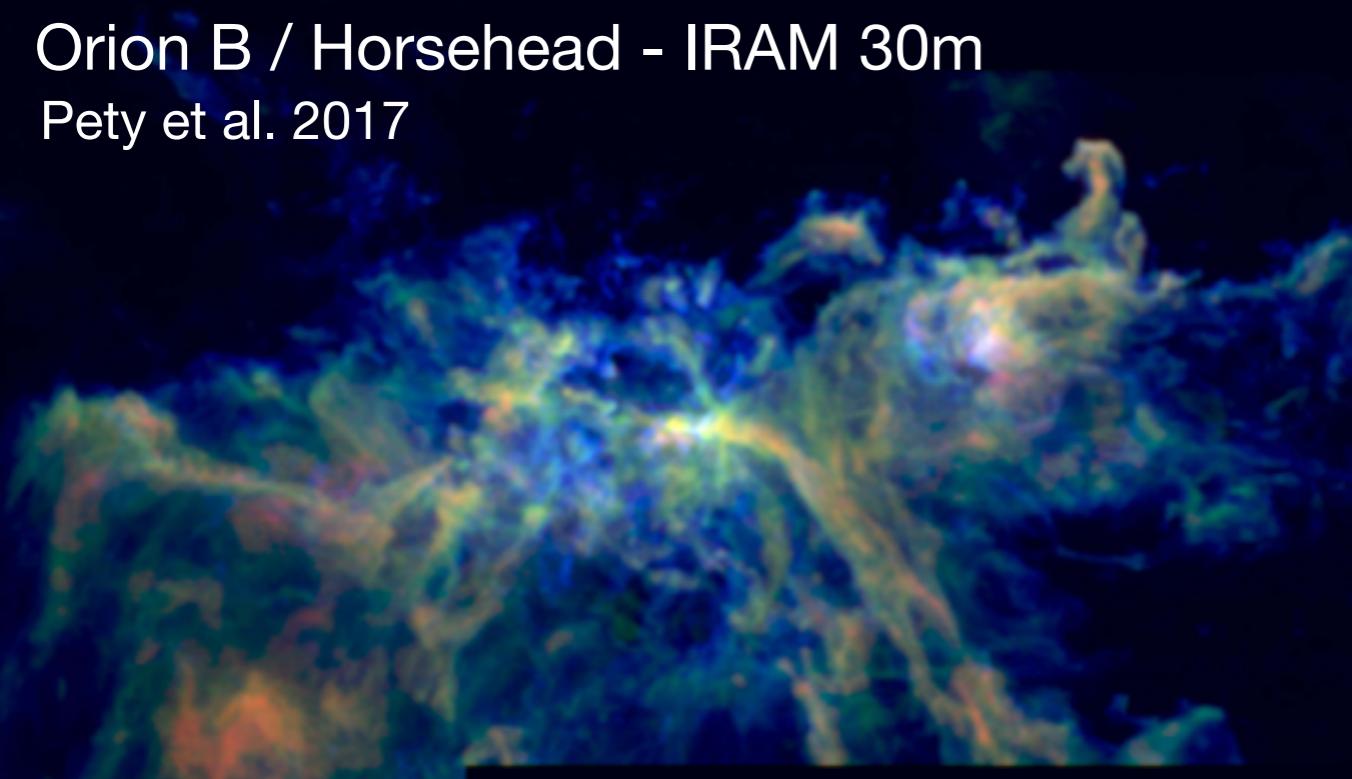
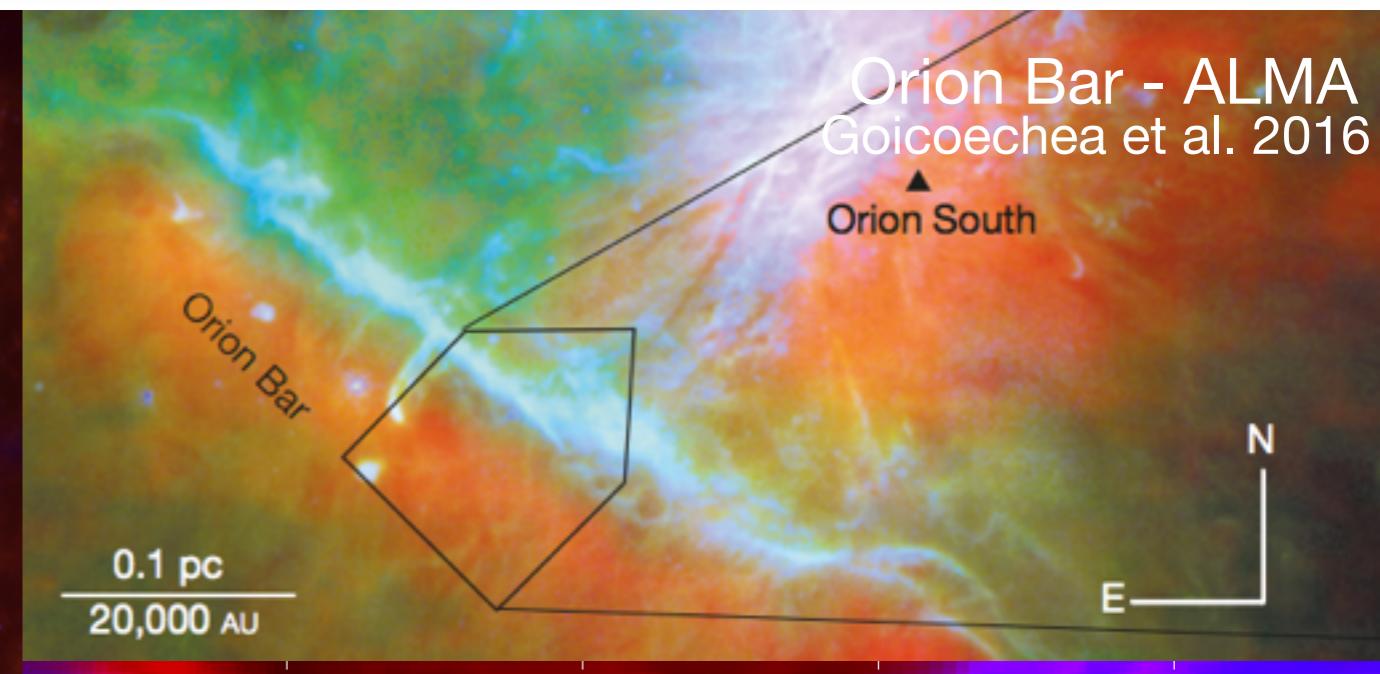
<http://ism.obspm.fr>

Franck Le Petit
David Languignon
Emeric Bron
Benjamin Godard

ISMDB

One of the services of the ANO5 "Plateforme MIS & Jets"

Goal: Provide services to prepare and interpret observations in Galactic & extragalactic interstellar medium



ISMDB

One of the services of the ANO5 "Plateforme MIS & Jets"

Goal: Provide services to prepare and interpret observations in the Galactic & extragalactic interstellar medium

Services are based on reference state-of-the-art codes

PDR code

Paris-Durham Shock code

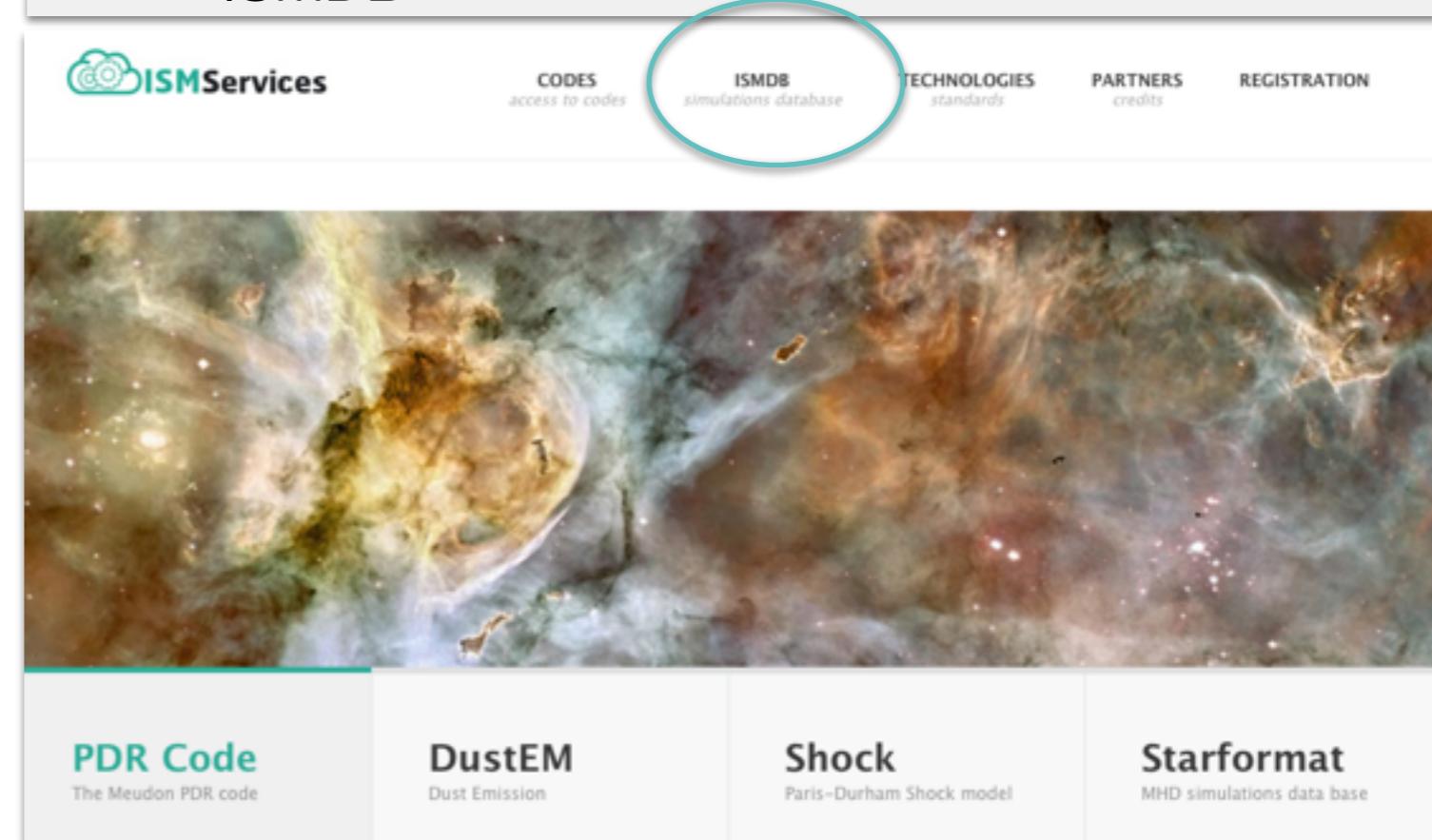
AMRVAC

TDR code

...

Several services are developed above the products of these codes

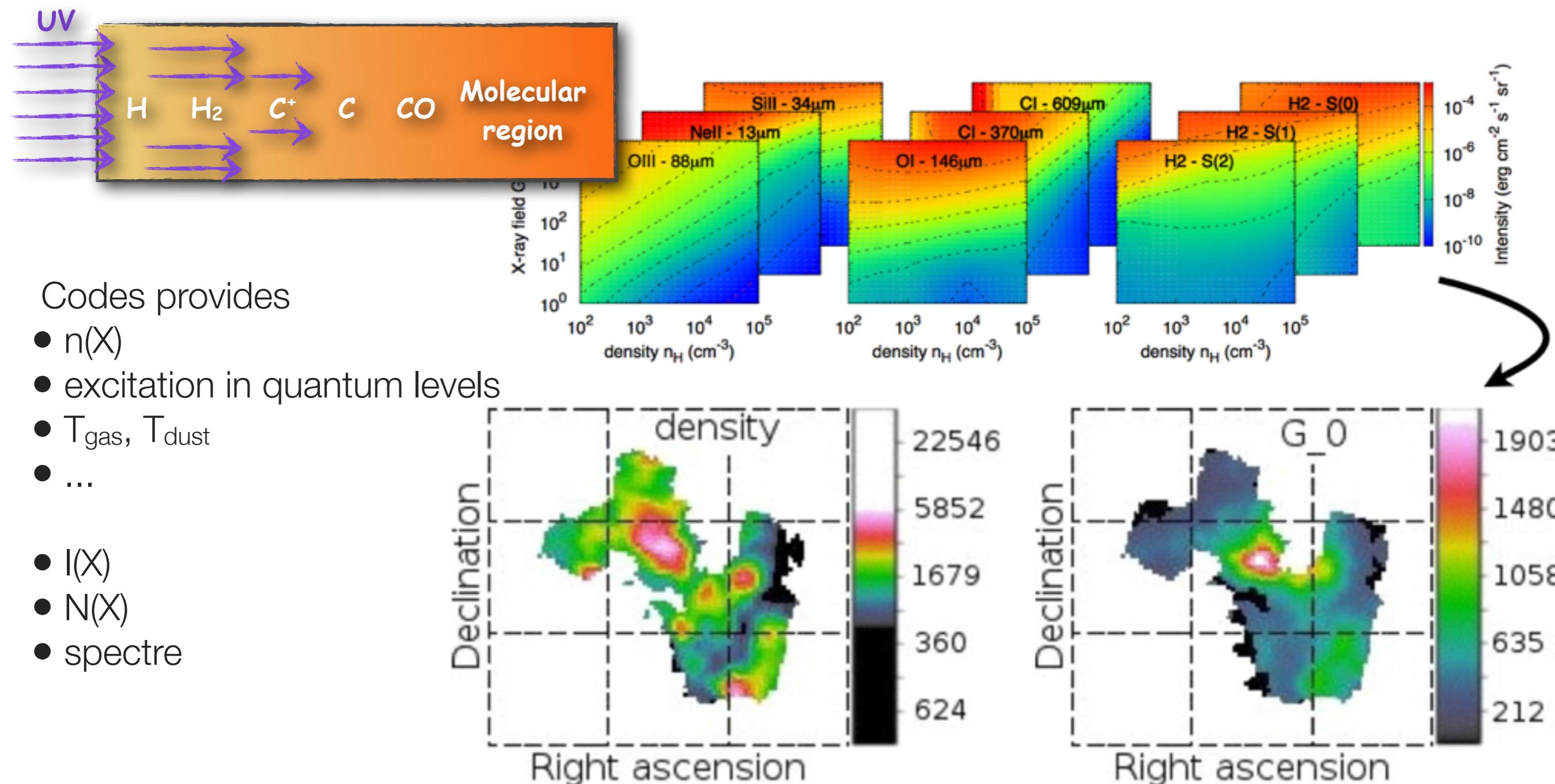
- Source codes & specific developments
- Online codes
- Tools to analyze results
 - Extractor & Chemistry Analyzer
- ISMDB



ISMDB

One of the services of the ANO5 "Plateforme MIS & Jets"

Goal: Provide services to prepare and interpret observations in Galactic & extragalactic interstellar medium

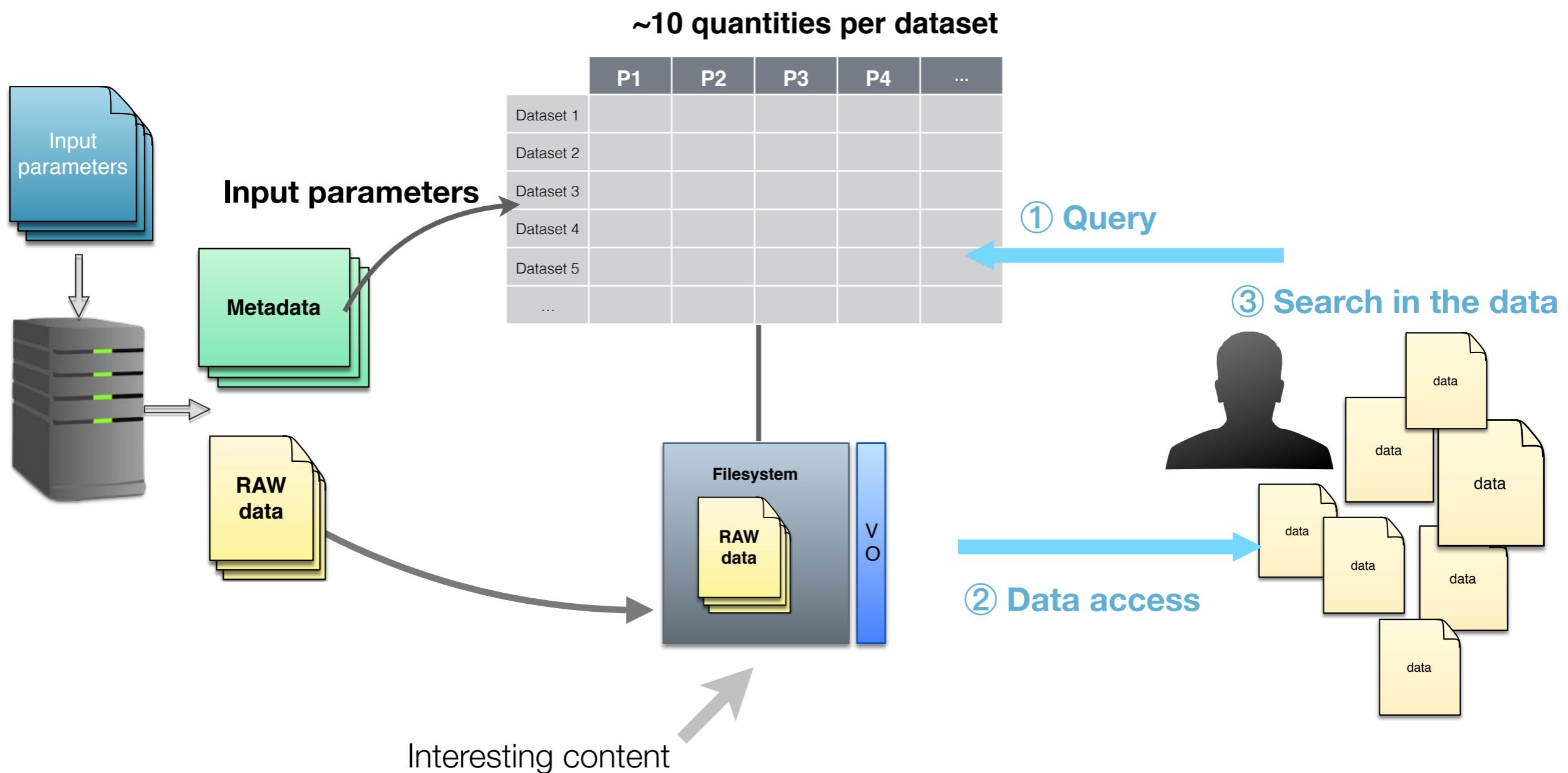


ISMDB

ISMDB: InterStellar Medium DataBase

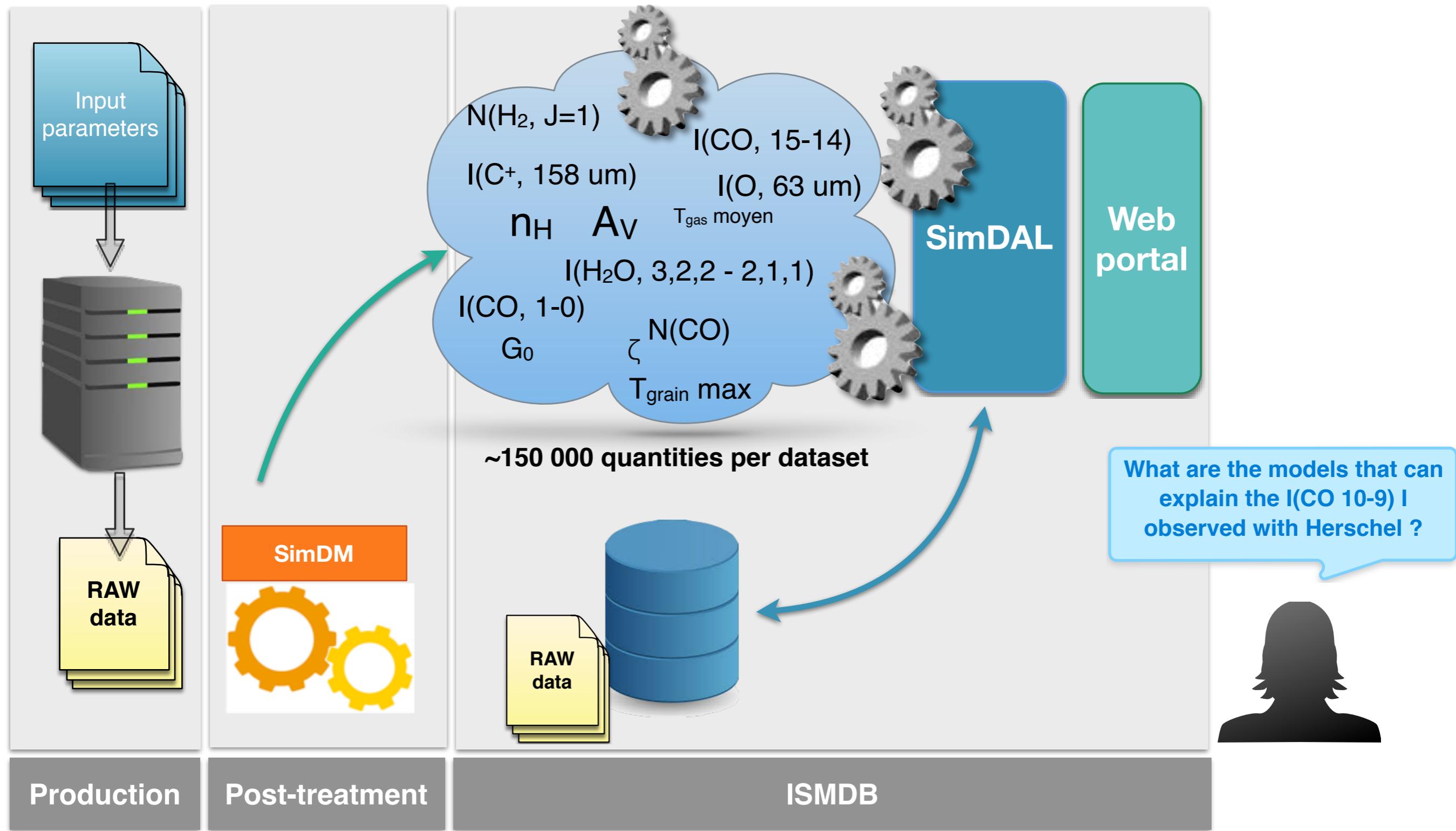
- not only a classical database to find pre-computed models
- but **also a tool that can *interpret* observations**

Standard databases:



ISMDB

- not only a classical database to find pre-computed models
- but also a tool that can *interpret* observations



ISMDB

Production of large grids of PDR models

- ~ 3000 PDR models in ISMDB for **standard galactic conditions**
- ~ 10 To of data

Cover: **PDR models for Herschel, IRAM/Noema, JWST, ALMA ... observations**

- Each model is characterized by + 150 000 queriable quantities
 - input parameters
 - computed **line intensities** (H_2 , CO, H_2O , ...)
 - **column densities** (total and in quantum levels)
- Raw data contains
 - all computed quantities (spatial profiles, ...)
 - **spectra**
 - ...

Models are stored in two projects

- isobaric models
- isochoric models

[Help](#) [Contact](#)

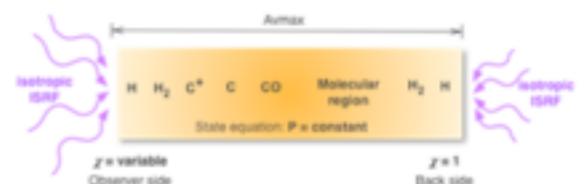
ISM DataBase – Inverse Search service Beta

Grid of isobaric PDR 1.5.2 models (2016-12-03)

code: PDR 1.5.2 rev 1714

Parameters

avmax: from 1.0 to 40.0 (mag)
radm_inl: from 1.0 to 1E5 (Mathis unit)
gas_pressure_input: from 1E5 to 1E9 (K cm⁻³)



A diagram illustrating an isobaric Photo-Dominated Region (PDR) model. It shows a central "Molecular region" containing chemical species like H, H₂, C⁺, C, and CO. Above the central region, it says "State equation: P = constant". The left side is labeled "Observer side" with "z = variable" below it. The right side is labeled "Back side" with "z = 1" below it. Wavy lines on either side represent "Isotropic ISRF" (Interstellar Radiation Field).

Description

This grid of isobaric PDR 1.5.2 models (revision 1714) covers photo-dominated regions conditions. Explored parameters are thermal pressure, UV field intensity and size of the clouds. The full grid contains 1372 2-side models where the back side of the cloud is submitted to the ISRF.

The chemistry takes into account 222 species, including C and O isotopes, linked by 6243 chemical reaction. No surface reactions are considered excepted for H₂. H₂ formation model takes into account Eley-Rideal and Langmuir-Hinshelwood mechanisms as described in Le Bourlot et al. (2012).

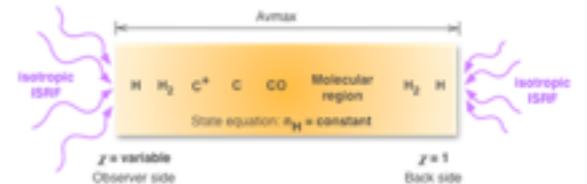
These models give access to all quantities computed by PDR 1.5.2 (line intensities, column densities, densities, temperature of gas and grains, ...).

Grid of isochoric PDR 1.5.2 models (2017-02-17)

code: PDR 1.5.2 rev 1787

Parameters

avmax: from 1.0 to 30.0 (mag)
proton_density_input: from 1E2 to 1E8 (cm⁻³)
radm_inl: from 1.0 to 1E5 (Mathis unit)



A diagram illustrating an isochoric Photo-Dominated Region (PDR) model. It shows a central "Molecular region" containing chemical species like H, H₂, C⁺, C, and CO. Above the central region, it says "State equation: n_H = constant". The left side is labeled "Observer side" with "z = variable" below it. The right side is labeled "Back side" with "z = 1" below it. Wavy lines on either side represent "Isotropic ISRF" (Interstellar Radiation Field).

Description

This grid of isochoric PDR 1.5.2 models (revision 1787) covers photo-dominated regions conditions. Explored parameters are proton density, UV field intensity and size of the clouds. The full grid contains 2128 2-side models where the back side of the cloud is submitted to the ISRF.

The chemistry takes into account 222 species, including C and O isotopes, linked by 6243 chemical reaction. No surface reactions are considered excepted for H₂. H₂ formation model takes into account Eley-Rideal and Langmuir-Hinshelwood mechanisms as described in Le Bourlot et al. (2012).

These models give access to all quantities computed by PDR 1.5.2 (line intensities, column densities, densities, temperature of gas and grains, ...).

ISM DataBase – Inverse Search service Beta

Grid of isobaric PDR 1.5.2 models

2016.12.03

1 – search among two parameters

x	Pgas_0	(cm ⁻³ _K)	<input checked="" type="checkbox"/> log scale
y	G0 observer side	(Mathis_unit)	<input checked="" type="checkbox"/> log scale

2 – fix all the other parameters

AVmax	(mag)	10
-------	-------	----

3 – observational constraints

Search for available quantities... Ex: N(H)	Use
---	-----

"I(CO v=0 J=1->v=0 J=0 angle 00 deg)" > 1.8E-7
"I(CO v=0 J=1->v=0 J=0 angle 00 deg)" < 2.4E-7
"I(H2 v=0 J=2->v=0 J=0 angle 60 deg)" > 1E-8
"I(H2 v=0 J=2->v=0 J=0 angle 60 deg)" < 5E-7

Search

① Select the searched input parameters

Example of a search:

- gas pressure
- UV intensity

② Fix the other input parameters

Example: size of the cloud

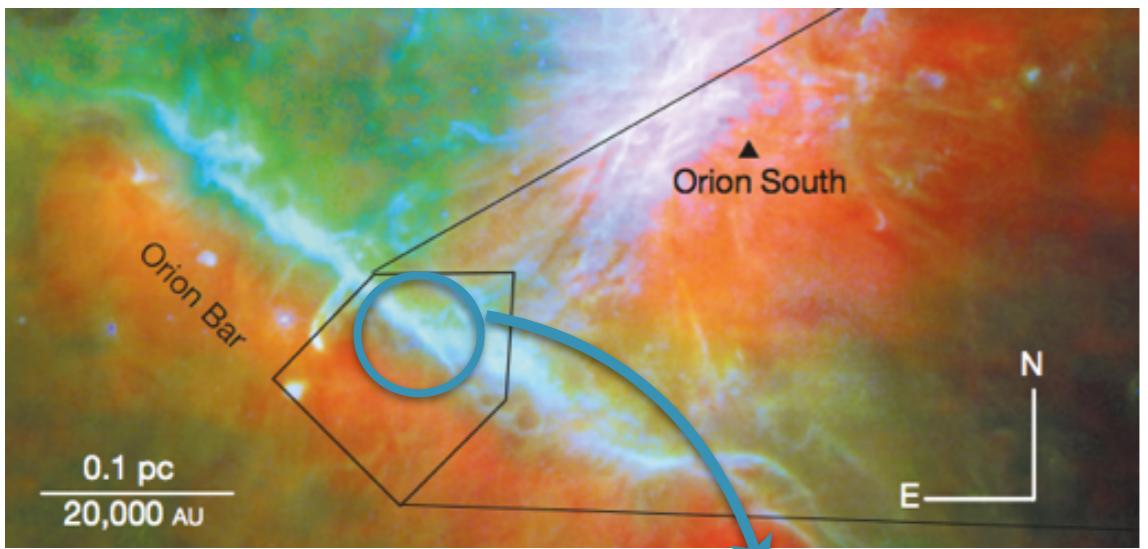
③ Enter the observations

Example: observations CO and H₂ intensities

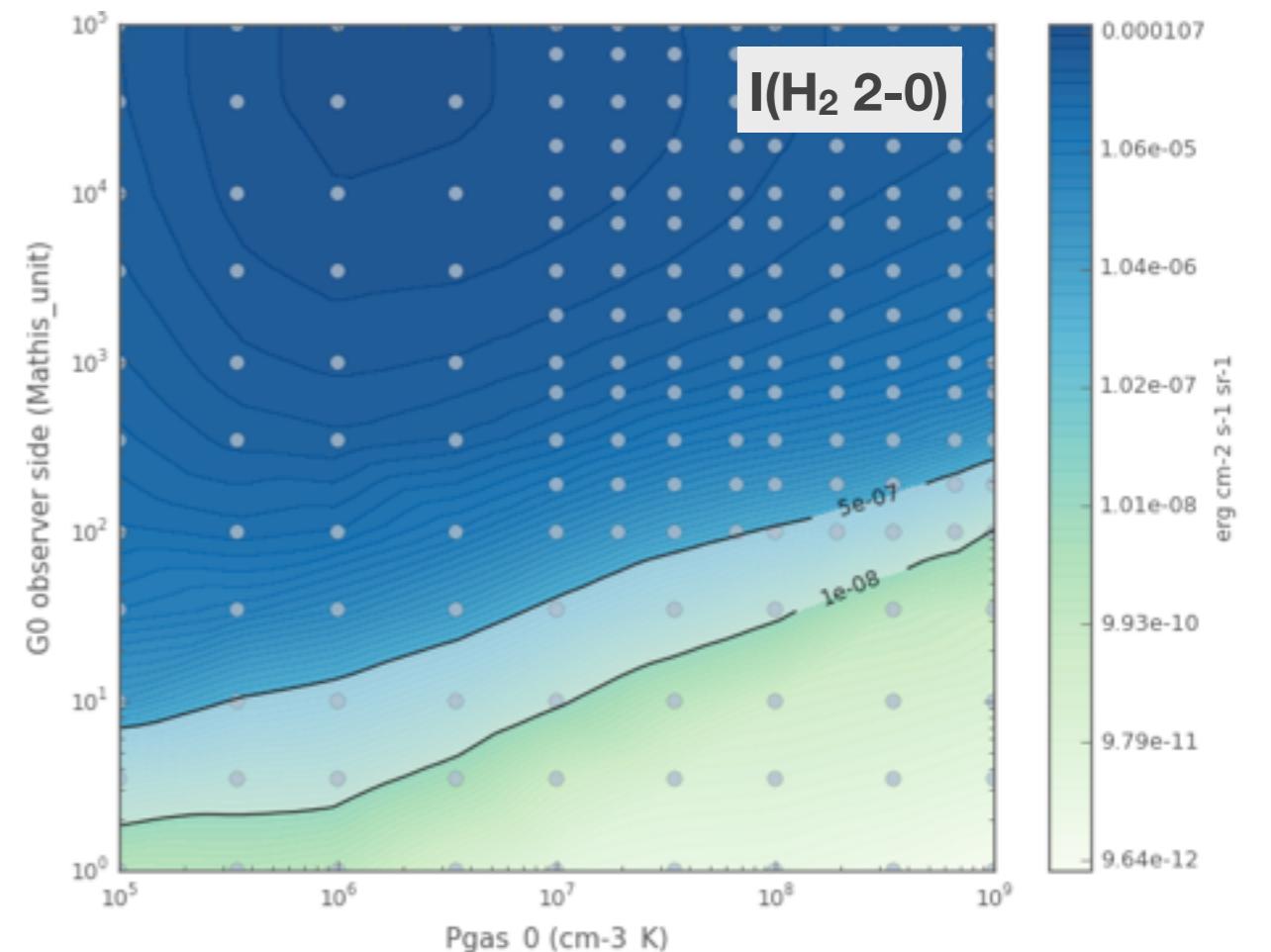
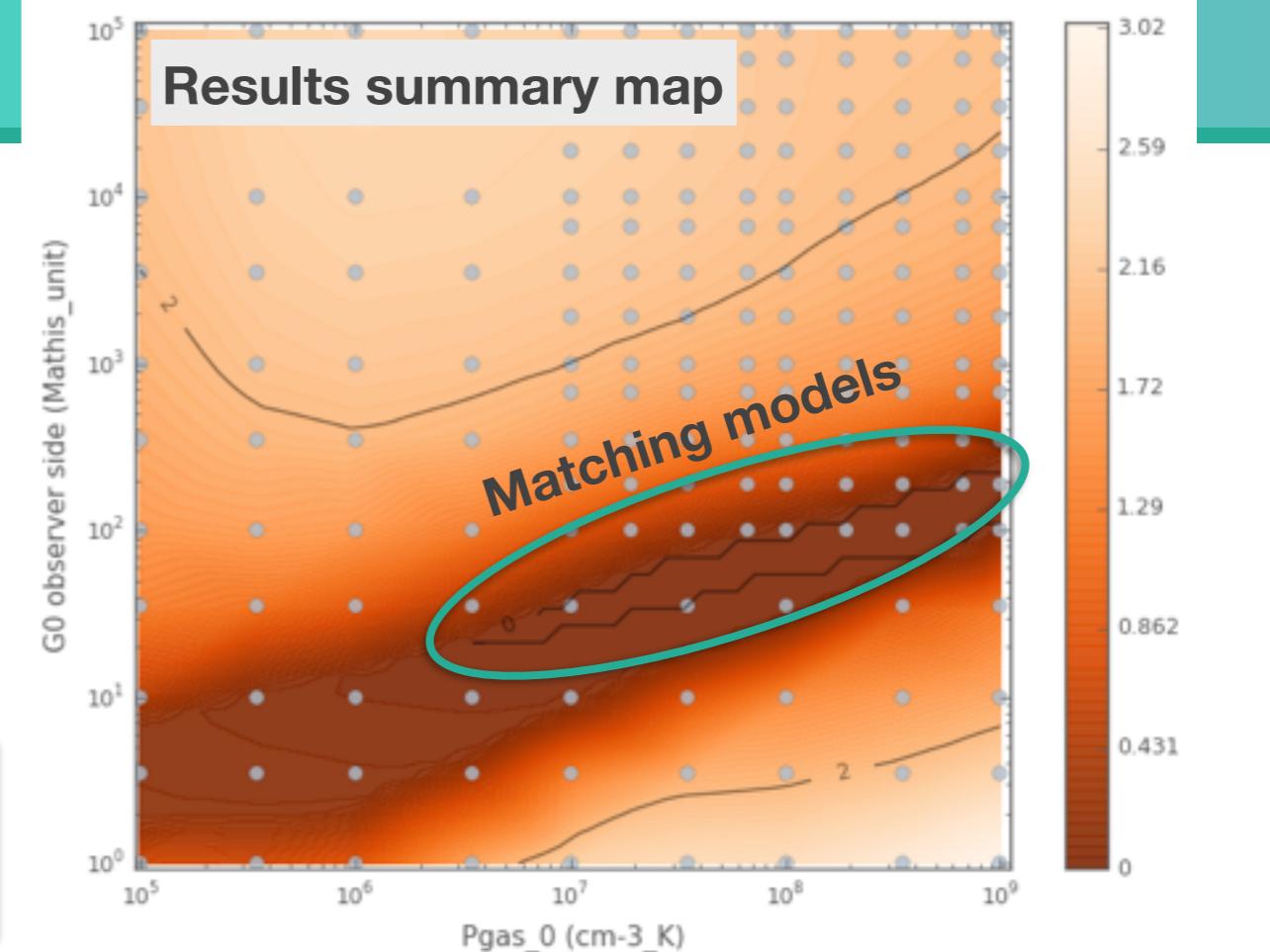
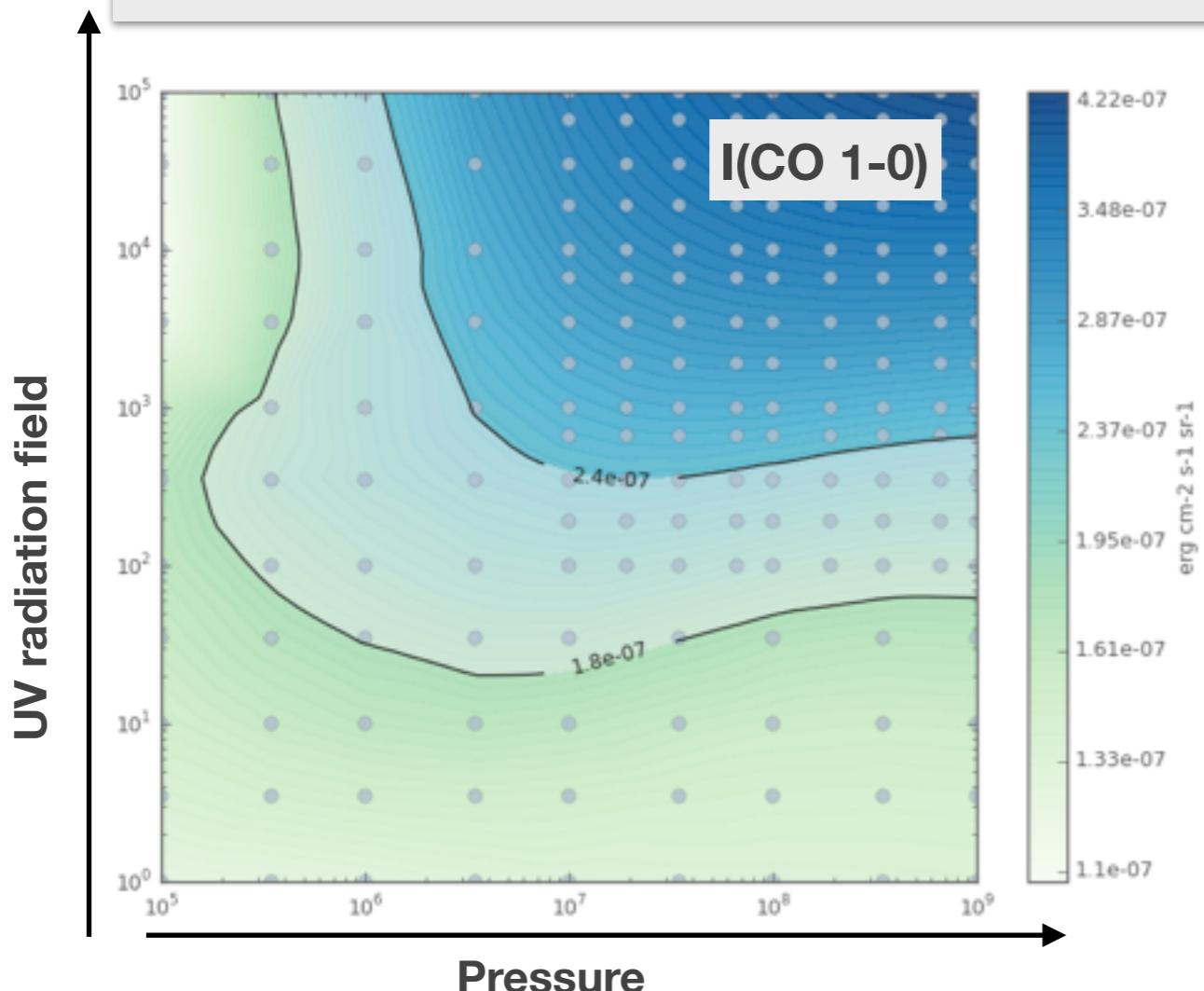
$1.8 \cdot 10^{-7} < I(\text{CO } 1-0) < 2.4 \cdot 10^{-7}$ erg cm⁻² s⁻¹ sr⁻¹

$1.0 \cdot 10^{-8} < I(\text{H}_2 \text{ } 2-0) < 5.0 \cdot 10^{-7}$ erg cm⁻² s⁻¹ sr⁻¹

Interpretation of observations



$$1.8 \cdot 10^{-7} < I(\text{CO } 1-0) < 2.4 \cdot 10^{-7} \text{ erg cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$
$$1.0 \cdot 10^{-8} < I(\text{H}_2 \text{ 2-0}) < 5.0 \cdot 10^{-7} \text{ erg cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$



Interpretation of observations

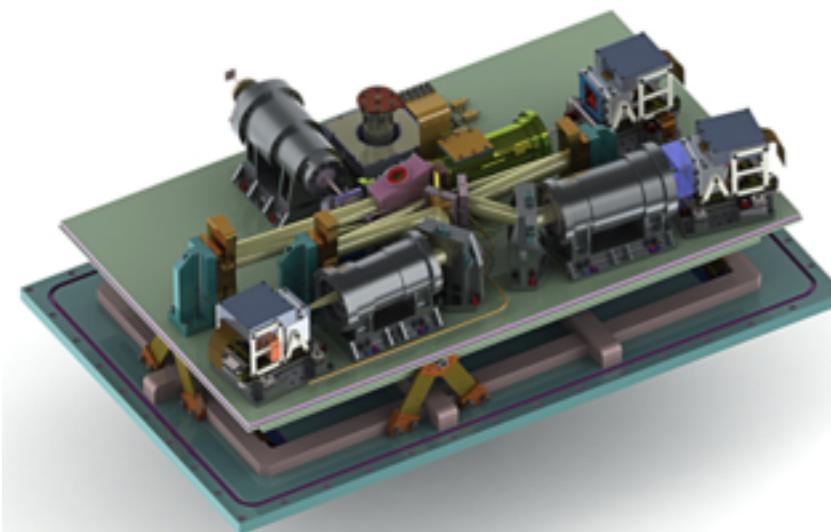
IGRINS observations

Instrument - Univ. Texas, Austin

Mc Donald observatory

Bands: H et K (1.5 to 2.5 microns)

R = 45 000

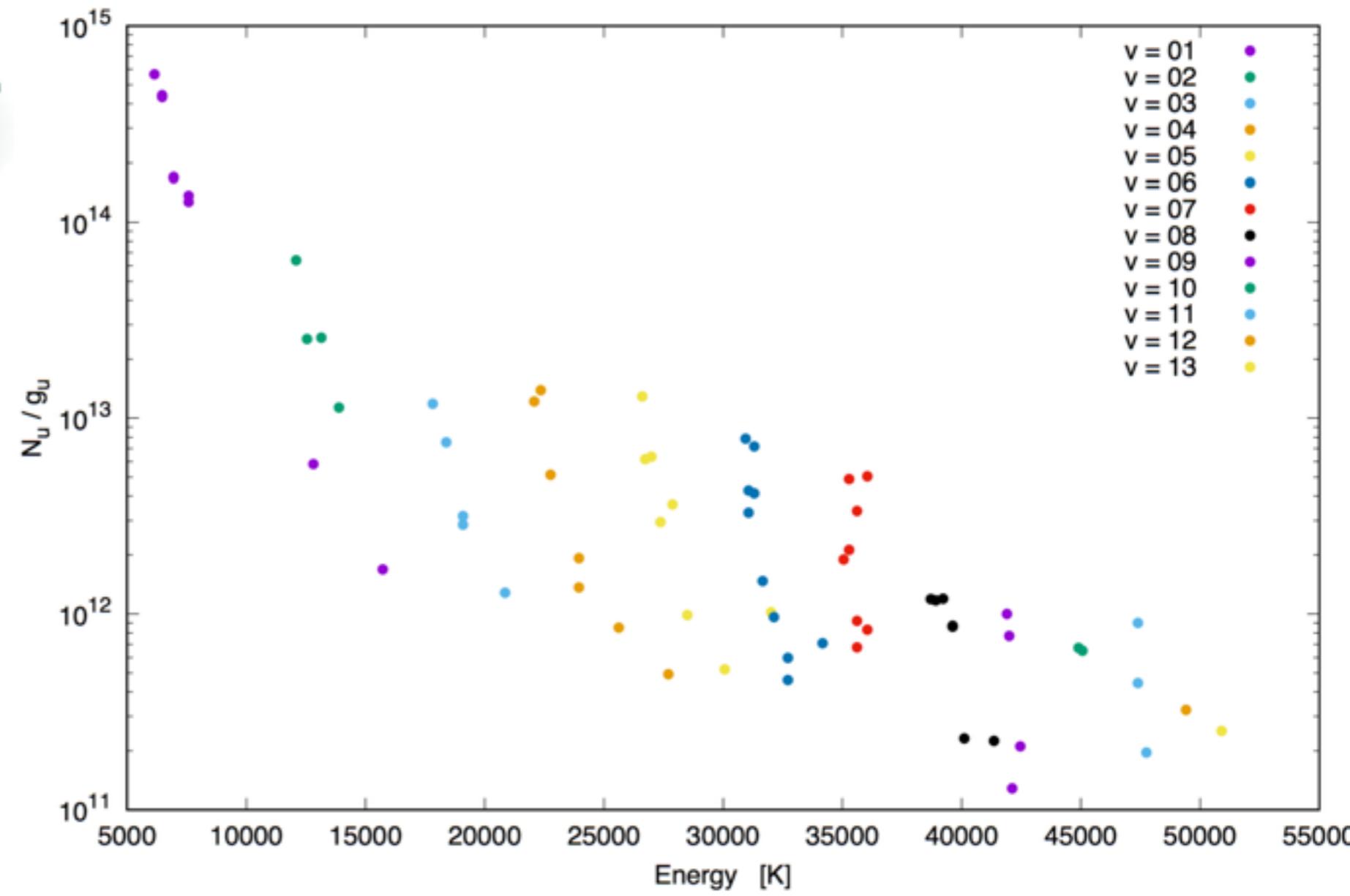


Observations of NGC 7023

(Le et al. - 2016 / ArXiV)

- Detection of **70 H₂ lines in NGC 7023**
- Conclude to a clumpy medium

H₂ excitation diagram at position A

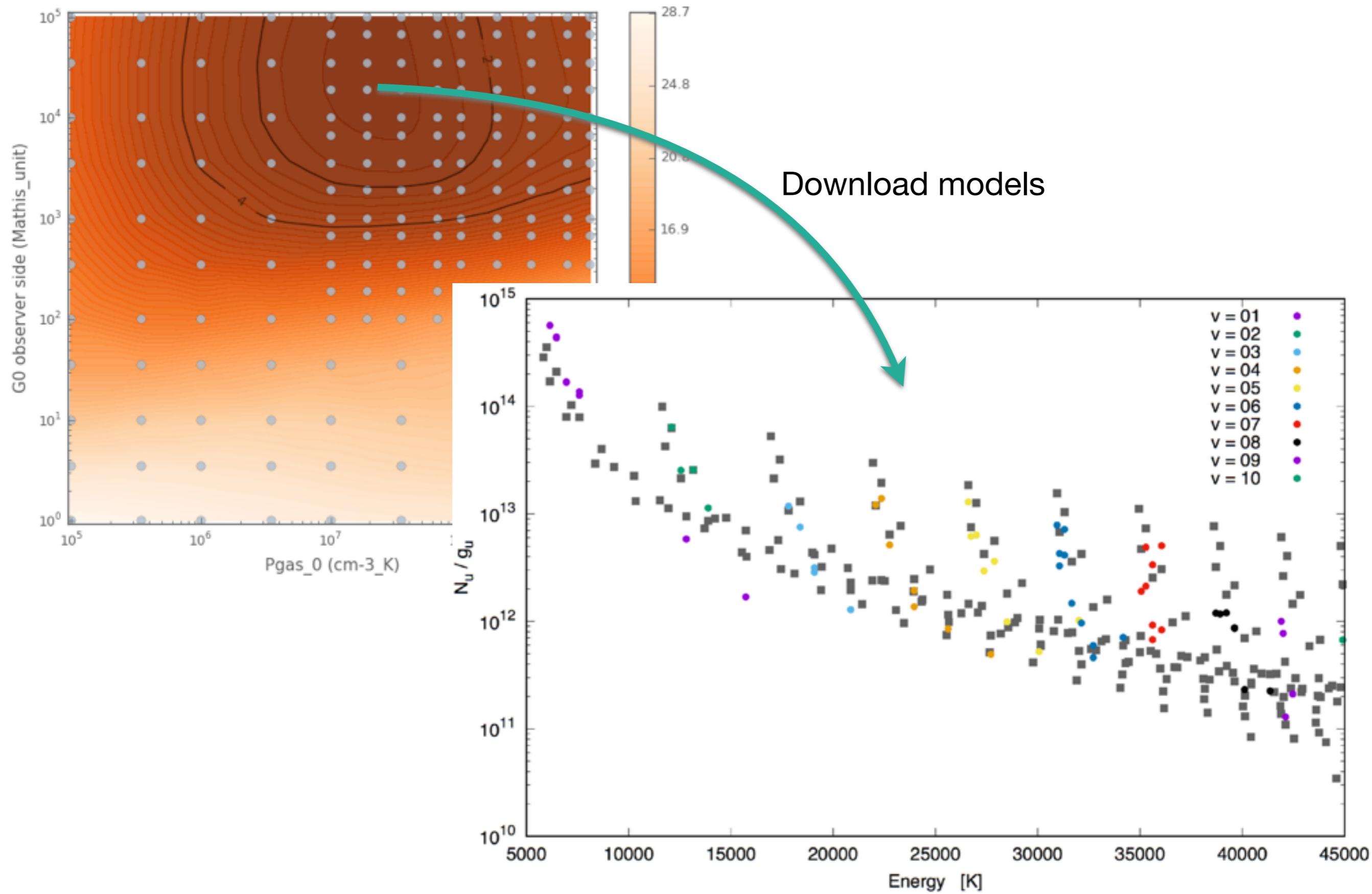


Interpretation of observations

Build the query for the 70 H₂ lines (140 constraints)

```
"I(H2 v=4,J=9->v=2,J=9 angle 60 deg)" < 5.408e-06
"I(H2 v=4,J=9->v=2,J=9 angle 60 deg)" > 2.912e-06
"I(H2 v=4,J=1->v=2,J=3 angle 60 deg)" < 2.665e-05
"I(H2 v=4,J=1->v=2,J=3 angle 60 deg)" > 1.435e-05
"I(H2 v=5,J=4->v=3,J=4 angle 60 deg)" < 9.659e-06
"I(H2 v=5,J=4->v=3,J=4 angle 60 deg)" > 5.201e-06
"I(H2 v=3,J=3->v=1,J=5 angle 60 deg)" < 1.2532e-05
"I(H2 v=3,J=3->v=1,J=5 angle 60 deg)" > 6.748e-06
"I(H2 v=5,J=5->v=3,J=5 angle 60 deg)" < 9.607e-06
"I(H2 v=5,J=5->v=3,J=5 angle 60 deg)" > 5.173e-06
"I(H2 v=6,J=2->v=4,J=0 angle 60 deg)" < 7.67e-06
"I(H2 v=6,J=2->v=4,J=0 angle 60 deg)" > 4.13e-06
"I(H2 v=10,J=1->v=7,J=3 angle 60 deg)" < 4.121e-06
"I(H2 v=10,J=1->v=7,J=3 angle 60 deg)" > 2.219e-06
"I(H2 v=5,J=0->v=3,J=2 angle 60 deg)" < 1.118e-05
"I(H2 v=5,J=0->v=3,J=2 angle 60 deg)" > 6.02e-06
"I(H2 v=5,J=7->v=3,J=7 angle 60 deg)" < 6.955e-06
"I(H2 v=5,J=7->v=3,J=7 angle 60 deg)" > 3.745e-06
"I(H2 v=4,J=2->v=2,J=4 angle 60 deg)" < 1.1531e-05
"I(H2 v=4,J=2->v=2,J=4 angle 60 deg)" > 6.209e-06
"I(H2 v=7,J=4->v=5,J=2 angle 60 deg)" < 5.109e-06
"I(H2 v=7,J=4->v=5,J=2 angle 60 deg)" > 2.751e-06
"I(H2 v=6,J=1->v=4,J=1 angle 60 deg)" < 1.846e-05
"I(H2 v=6,J=1->v=4,J=1 angle 60 deg)" > 9.94e-06
"I(H2 v=6,J=2->v=4,J=2 angle 60 deg)" < 1.599e-05
"I(H2 v=6,J=2->v=4,J=2 angle 60 deg)" > 8.61e-06
"I(H2 v=5,J=9->v=3,J=9 angle 60 deg)" < 1.729e-05
"I(H2 v=5,J=9->v=3,J=9 angle 60 deg)" > 9.31e-06
"I(H2 v=5,J=1->v=3,J=3 angle 60 deg)" < 2.379e-05
"I(H2 v=5,J=1->v=3,J=3 angle 60 deg)" > 1.281e-05
"I(H2 v=13,J=1->v=9,J=1 angle 60 deg)" < 9.334e-07
"I(H2 v=13,J=1->v=9,J=1 angle 60 deg)" > 5.026e-07
"I(H2 v=6,J=3->v=4,J=3 angle 60 deg)" < 1.287e-05
"I(H2 v=6,J=3->v=4,J=3 angle 60 deg)" > 6.93e-06
I(H2 v=7,J=3->v=5,J=1 angle 60 deg)" < 1.1531e-05
I(H2 v=7,J=3->v=5,J=1 angle 60 deg)" > 6.209e-06
I(H2 v=4,J=3->v=2,J=5 angle 60 deg)" < 1.2961e-05
I(H2 v=4,J=3->v=2,J=5 angle 60 deg)" > 6.979e-06
I(H2 v=6,J=4->v=4,J=4 angle 60 deg)" < 3.523e-06
I(H2 v=6,J=4->v=4,J=4 angle 60 deg)" > 1.897e-06
I(H2 v=6,J=5->v=4,J=5 angle 60 deg)" < 7.878e-06
I(H2 v=6,J=5->v=4,J=5 angle 60 deg)" > 4.242e-06
I(H2 v=3,J=5->v=1,J=7 angle 60 deg)" < 2.457e-06
I(H2 v=3,J=5->v=1,J=7 angle 60 deg)" > 1.323e-06
I(H2 v=11,J=1->v=8,J=1 angle 60 deg)" < 2.899e-06
I(H2 v=11,J=1->v=8,J=1 angle 60 deg)" > 1.561e-06
I(H2 v=7,J=2->v=5,J=0 angle 60 deg)" < 4.628e-06
I(H2 v=7,J=2->v=5,J=0 angle 60 deg)" > 2.492e-06
I(H2 v=8,J=7->v=6,J=5 angle 60 deg)" < 5.824e-06
I(H2 v=8,J=7->v=6,J=5 angle 60 deg)" > 3.136e-06
I(H2 v=5,J=2->v=3,J=4 angle 60 deg)" < 9.425e-06
I(H2 v=5,J=2->v=3,J=4 angle 60 deg)" > 5.075e-06
I(H2 v=6,J=0->v=4,J=2 angle 60 deg)" < 9.776e-06
I(H2 v=6,J=0->v=4,J=2 angle 60 deg)" > 5.264e-06
I(H2 v=6,J=7->v=4,J=7 angle 60 deg)" < 1.2532e-05
I(H2 v=6,J=7->v=4,J=7 angle 60 deg)" > 6.748e-06
I(H2 v=11,J=3->v=8,J=3 angle 60 deg)" < 1.924e-06
I(H2 v=11,J=3->v=8,J=3 angle 60 deg)" > 1.036e-06
I(H2 v=1,J=11->v=0,J=9 angle 60 deg)" < 4.407e-06
I(H2 v=1,J=11->v=0,J=9 angle 60 deg)" > 2.373e-06
I(H2 v=8,J=5->v=6,J=3 angle 60 deg)" < 5.122e-06
I(H2 v=8,J=5->v=6,J=3 angle 60 deg)" > 2.758e-06
I(H2 v=7,J=1->v=5,J=1 angle 60 deg)" < 1.2922e-05
I(H2 v=7,J=1->v=5,J=1 angle 60 deg)" > 6.958e-06
I(H2 v=8,J=4->v=6,J=2 angle 60 deg)" < 5.109e-06
I(H2 v=8,J=4->v=6,J=2 angle 60 deg)" > 2.751e-06
I(H2 v=6,J=1->v=4,J=3 angle 60 deg)" < 2.405e-05
...
...
...
...
```

Interpretation of observations



VO Integration

Metadata definition & organization

Data Access

Raw data (extractor tool)

Extractor Tool

PDR Extractor

n(CO) Confirm Remove All

DM54NoPAH_A1e1p6p7e7r1e4_s_20.hdf5

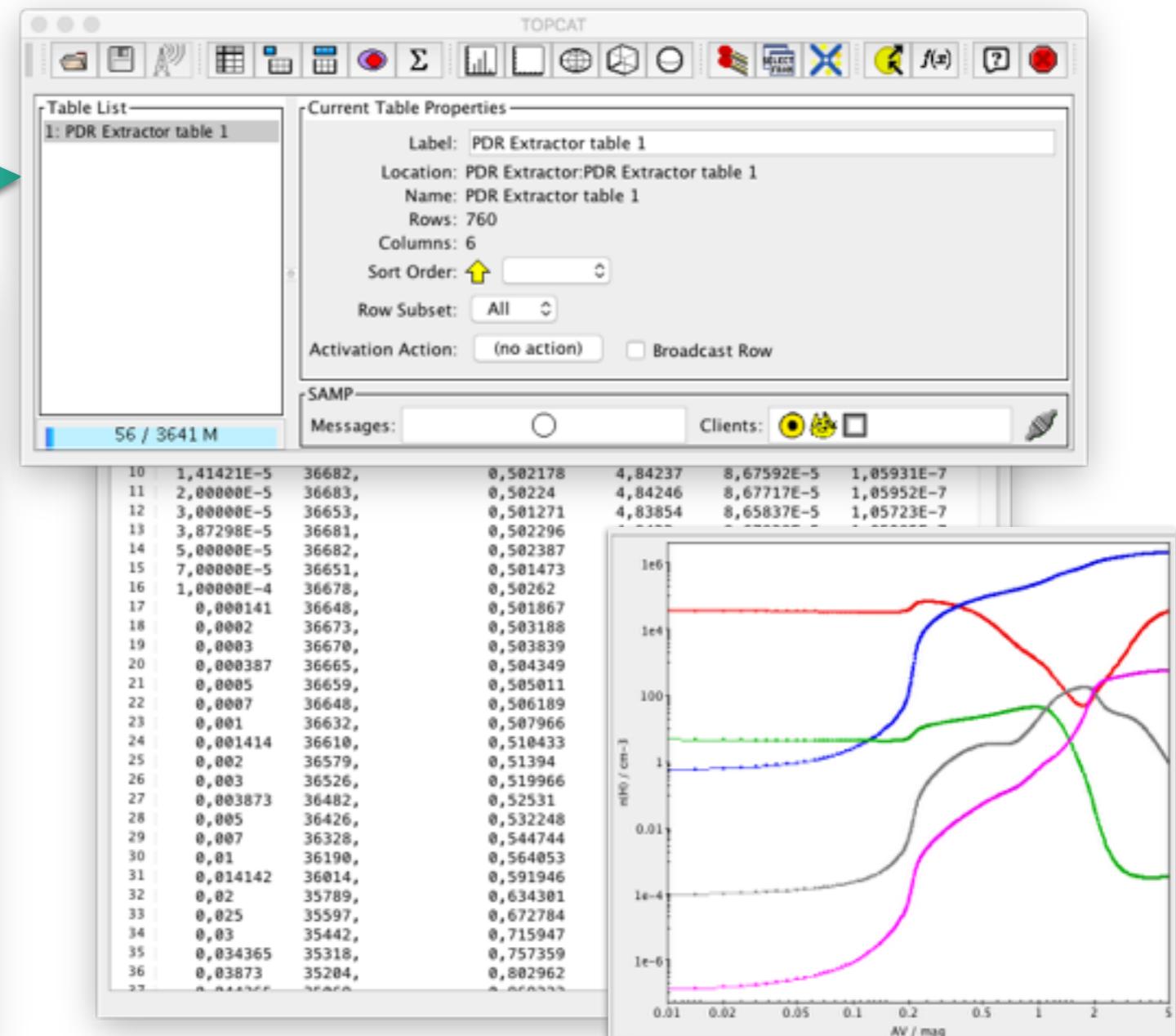
- Integrated quantities
- Local quantities
 - Auxiliary
 - Excitation
 - H₂ chemistry
 - Ice layers
 - Molecular fraction
 - Photo reactions
 - Radiation
 - Thermal balance
 - Densities
 - Column densities
 - Densities
 - Dust
 - Gas state
 - Positions
 - AV
 - Distance
 - tauV
 - Parameters

Export as Text Export as VOTable Send Table

Simulation Data Model (SimDM)

Simulation Data Access Layer (SimDAL)

VOTable
SAMP connector



Technical challenges

Large amount of metadata → 2 main difficulties

1 - high dimension database

Standard solution: RDBMS / SQL

	Meta1	Meta2	Meta3	Meta4	...
Dataset 1					
Dataset 2					
Dataset 3					
Dataset 4					
Dataset 5					
...					

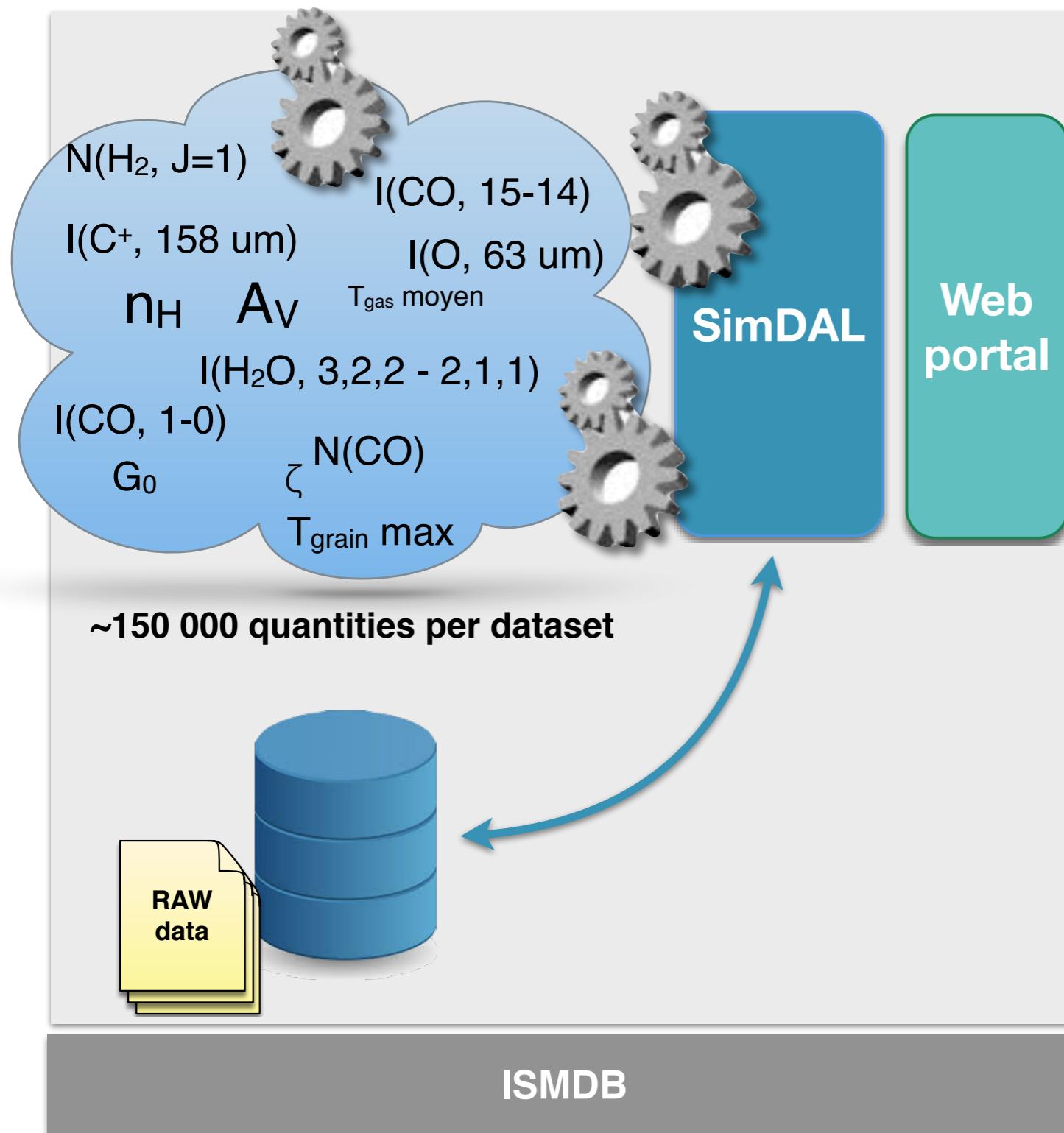
	Number of col.
MySQL	4096
Postgress	250 - 1600
Oracle	1000
Microsoft	30000

Standard solutions for databases (RDBMS/SQL) are not adapted to high dimensions / heterogeneous data

ISMDB must manipulate more than 100 000 dimensions

Technical challenges

2 - interaction between a human and the system



What are the models that can explain the $I(CO\ 10-9)$ I observed with Herschel ?



How the user can know the name of the available quantities ?

Technical challenges

VLA archive Interface

NRAO Science Data Archive : Advanced Search Tool
Historical VLA, Jansky VLA, VLBA and GBT Data Products

Output Control Parameters :

Choose Query Return Type :

Download Archive Data Files
 VLA Observations Summary
 List of Observation Scans
 List of Projects

General Search Parameters :

Telescopes All Jansky VLA Historical VLA VLBA GBT

Project Code Project Session Dates From
GBT: AGBT12A_055
JVLA: 12A-256

Observer Name Archive File ID
(partial strings allowed) To
(2010-06-21 14:20:30)

Position Search :

Target Name Search Type Min. Exposure (secs)
RA or Longitude (04h33m11.1s or 68.29d) DEC or Latitude (05d21'15.5" or 5.352d) Equinox

Search Radius 1.0' (1d00'00" or 0.2d) - OR - Check for automatic VLA field-of-view, freq. dependent.??

Observing Configurations Search :

Telescope All A AB BnA B BC CnB
Config C CD DnC D DA

Sub_array All 1 2 3 4 5

Polarization Data Type

Observing Bands All 4 P L S C
 X U K Ka Q W

Frequency Range (In MHz : 1665.401 - 1720.500)

Enter Locked Project Access key : Unique keywords may be used to unlock proprietary data from individual observing projects. Contact the [NRAO Data Analysts](#) for project access keys.

23 paramètres de recherche
Interface complexe

Technical challenges

VLA archive Interface

NRAO Science Data Archive : Advanced Search Tool
Historical VLA, Jansky VLA, VLBA and GBT Data Products

Output Control Parameters :

Choose Query Return Type :

Download Archive Data Files
 VLA Observations Summary
 List of Observation Scans
 List of Projects

Output Tbl Format Sort Order Column 1
Max Output Tbl Rows Sort Order Column 2

General Search Parameters :

Telescopes All Jansky VLA Historical VLA VLBA GBT

Project Code Project Session Dates From
GBT: AGBT12A_055
JVLA: 12A-256

Observer Name Archive File ID
(partial strings allowed) To

Position Search :

Target Name Search Type Min. Exposure (secs)
RA or Longitude DEC or Latitude Equinox

Search Radius - OR - Check for automatic VLA field-of-view, freq. dependent.??

Observing Configurations Search :

Telescope All A AB BnA B BC CnB
Config C CD DnC D DA

Sub_array All 1 2 3 4 5

Polarization Data Type

Observing Bands All 4 P L S C
 X U K Ka Q W

Frequency Range (In MHz : 1665.401 - 1720.500)

Enter Locked Project Access key :
Unique keywords may be used to unlock proprietary data from individual observing projects. Contact the [NRAO Data Analysts](#) for project access keys.

23 paramètres de recherche
Interface complexe

ISMDB
150 000 parameters !

Semantics

ISM Services CODES ISMDB

ISM DataBase – Inverse Search service Beta

Grid of isobaric PDR 1.5.2 models
2016.12.03

1 – search among two parameters

x Pg_{as}_0 (cm⁻³_K) log scale

y G₀ observer side (Mathis_unit) log scale

2 – fix all the other parameters

AVmax (mag) 10

3 – observational constraints

Search for available quantities... Ex: N(H) Use

```
"I(CO v=0,J=1->v=0,J=0 angle 00 deg)" > 1.8E-7
"I(CO v=0,J=1->v=0,J=0 angle 00 deg)" < 2.4E-7
"I(H2 v=0,J=2->v=0,J=0 angle 60 deg)" > 1E-8
"I(H2 v=0,J=2->v=0,J=0 angle 60 deg)" < 5E-7
```

Search

3 – observational constraints

Search for available quantities... Ex: N(H)

Use

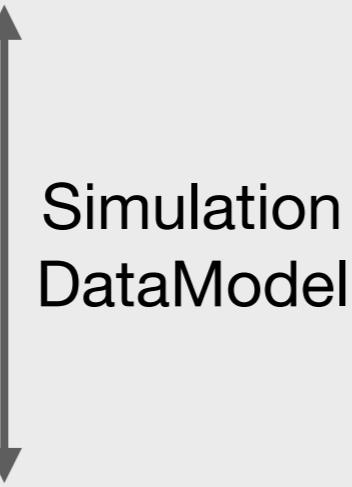


Semantics interpreter

Semantics

Each metadata is taggued by:

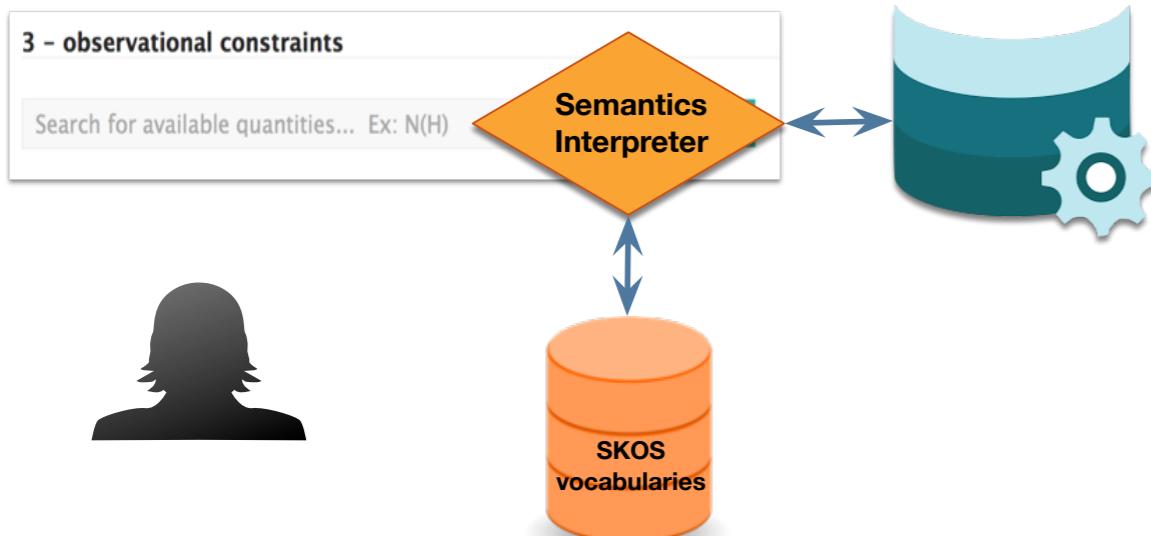
- ID
- name
- unit
- utype
- description
- **label (UCD / SKOS)**
- ...



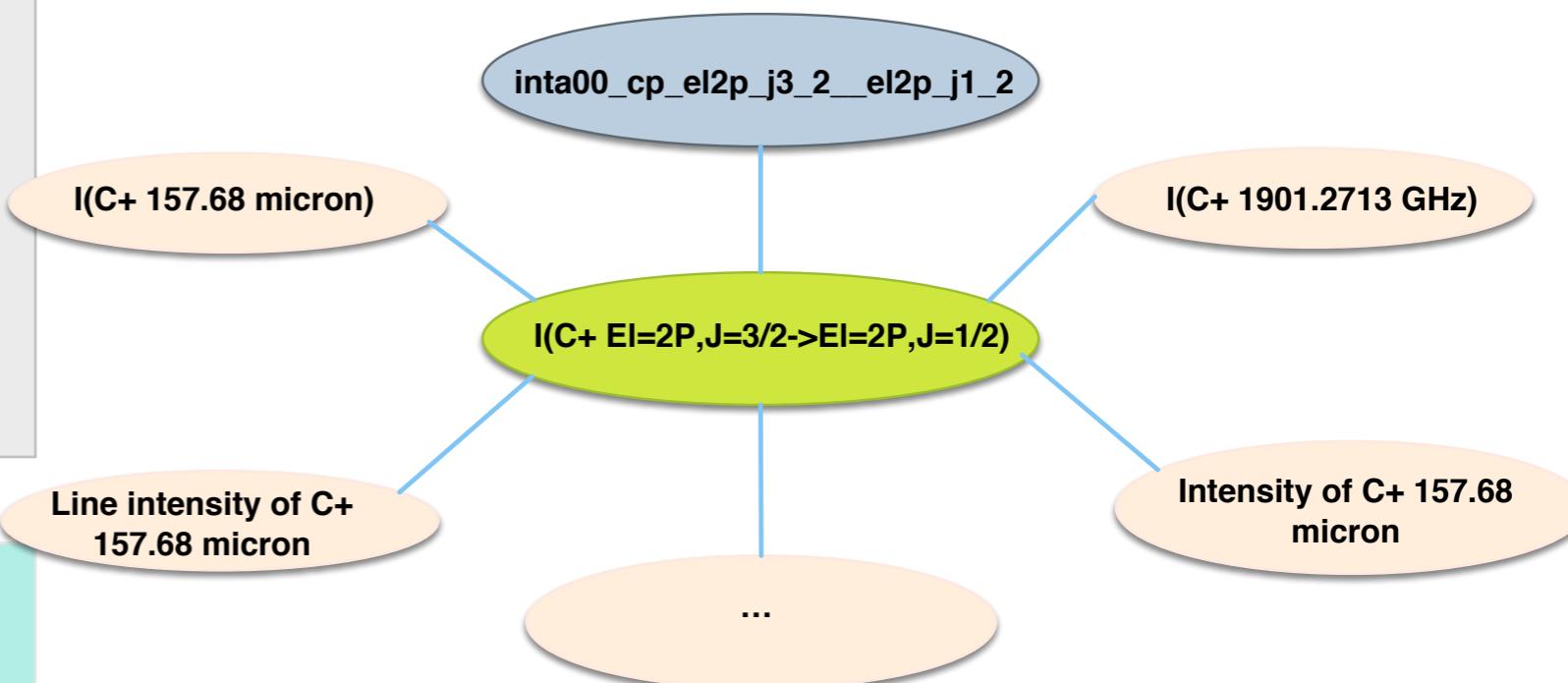
SKOS vocabulary

For each quantity several synonyms
(name, units, ...)

~ 300 000 terms for the PDR code



Example of the 157.7 micron C+ line intensity

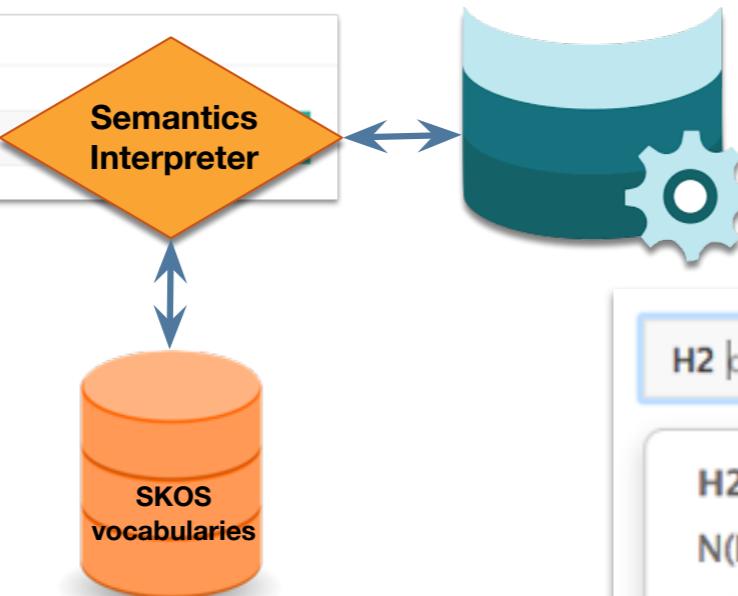


ID: inta00_cp_el2p_j3_2_el2p_j1_2
PREF: I(C+ El=2P,J=3/2->El=2P,J=1/2)
ALT: I(C+ El=2P,J=3/2->El=2P,J=1/2) face on
ALT: I(C+ 157.68 micron) face on
ALT: Intensity of C+ 157.68 micron face on
ALT: Line intensity of C+ 157.68 micron face on
ALT: I(C+ 1901.2713 GHz) face on
ALT: Intensity of C+ 1901.2713 GHz face on
ALT: Line intensity of C+ 1901.2713 GHz face on
...

Technical challenges

3 - observational constraints

Search for available quantities... Ex: N(H)



Semantics Interpreter

Semantics

SKOS: PREF + ALT
→ synonyms

+

Ranking system

(learn from users)

H₂ column density

Use

H₂ column density

N(H₂)

N(C₂H₂)

N(c-C₃H₂)

N(C₁₃CH₂)

N(C₁₃CH₂ +)

C₂H₂ column density

Column density of H₂

I(H₂ | 0-0 S(0)) angle 00 degrees

Use

I(H₂ 0-0 S(0)) angle 00 degrees

I(H₂ 10-10 S(0)) angle 00 degrees

I(H₂ 9.6645 micrometres) angle 00 degrees

I(H₂ 28.2196 micrometres) angle 00 degrees

I(H₂ 156.4883 micrometres) angle 00 degrees

I(H₂ v=0,J=2->v=0,J=0) angle 00 degrees

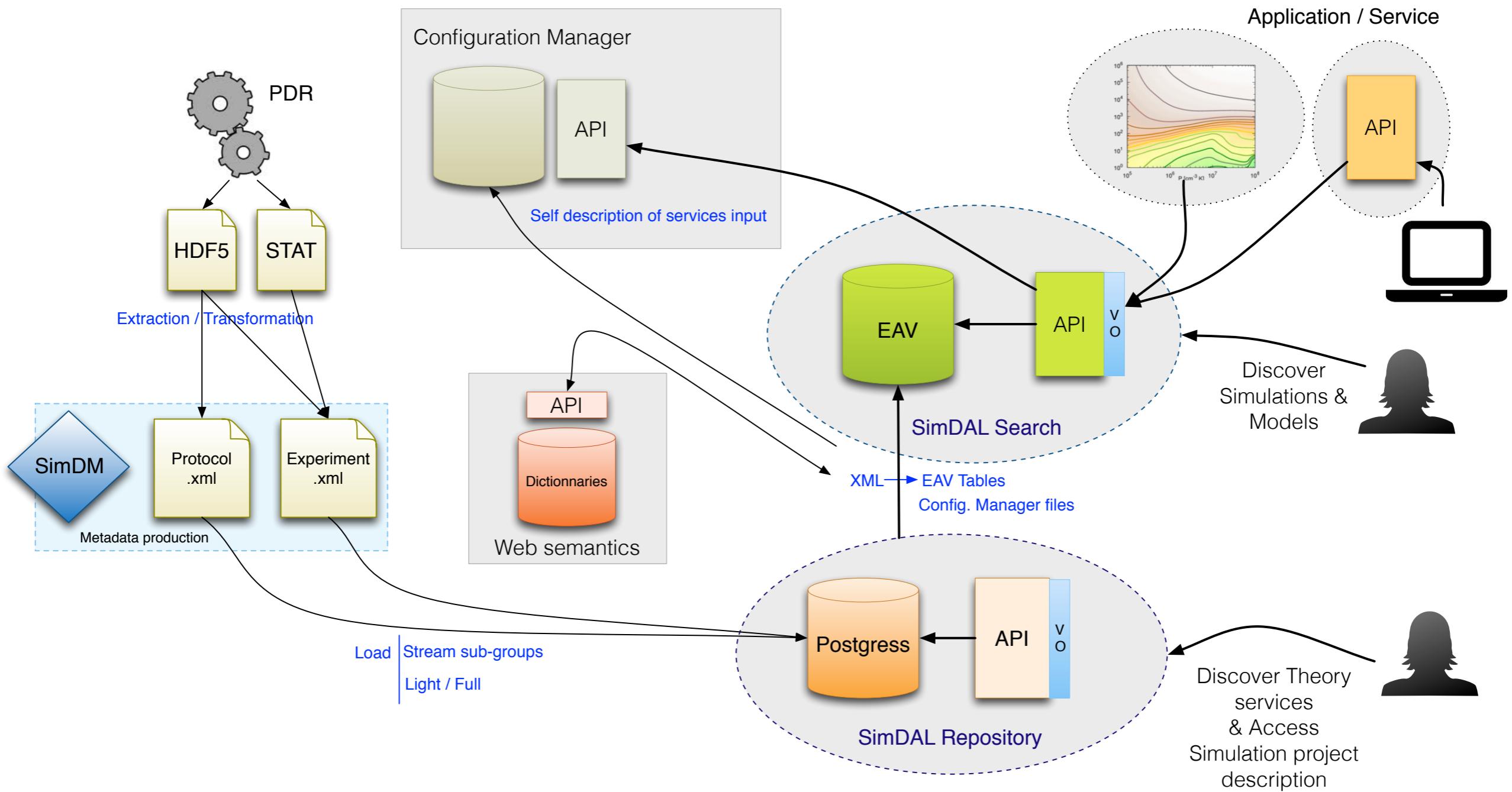
I(H₂O 6.1140 cm⁻¹) angle 00 degrees

I(H₂O J=1,ka=1,kc=1->J=0,ka=0,kc=0) angle 00 degrees

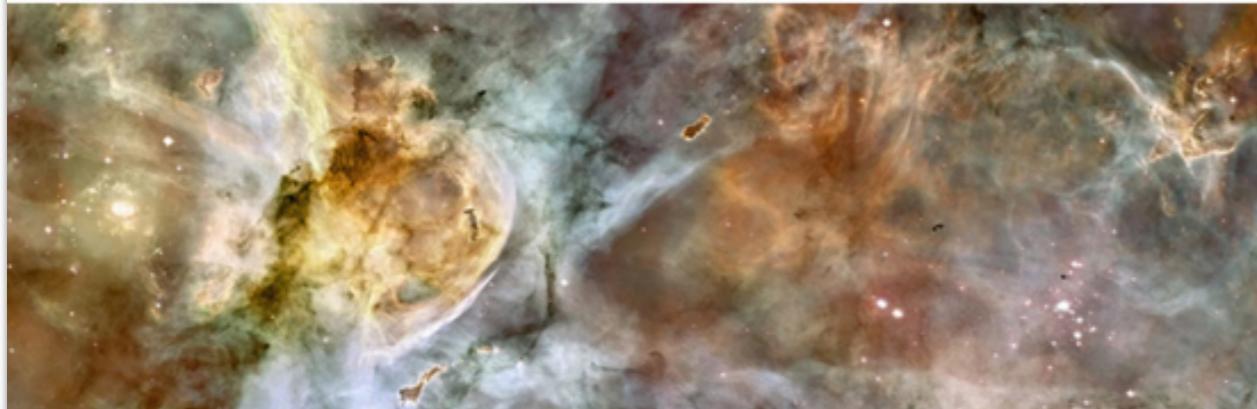
ISMDB infrastructure

ISMDB infrastructure development

- **modular**
- based on **robust & mature technologies**
- **generic**: can integrate models from any similar code than the PDR code



ISMDB



The screenshot shows the ISM Services website. At the top, there's a navigation bar with links for ISM Services, CODES (access to codes), ISMDB (simulations database), TECHNOLOGIES (standards), PARTNERS (credits), and REGISTRATION. Below the navigation is a large image of a star-forming region in the Carina Nebula. Underneath the image are four cards: PDR Code (The Meudon PDR code), DustEM (Dust Emission), Shock (Paris-Durham Shock model), and Starformat (MHD simulations data base). At the bottom of the page is the ISM DataBase – Inverse Search service (Beta) section, which includes input fields for parameters like Pgash_0 and G0 observer side, and a search interface for observational constraints.

Status

- Public at <http://ism.obspm.fr>
- Grids of PDR models

Starts to be used:

- Individual teams
- Projects as SPICA, GUSTO (NASA/CNES)
- JWST ERS

Plans - short term

- Semantics
- Grids of shocks models
- Operations on quantities

Plans - medium term

- Other ways to query models
- Quickviews on models

Plans - long term

- Search in N-dimension space
- Interpretation at each pixels of maps

VO-Theory

Status on standards



Franck Le Petit
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VO-Theory standards

SimDM

- approved a few years ago
- no evolution

SimDAL

Access to theoretical data
(consider big / heterogeneous data)

- 2 reference implementations + client
- all WG / IG have done their comments
- few modifications (clarifications of the text)
- start of implementations on other services

Everything ready for the integration
of theoretical data in the Virtual
Observatory

Both documents on the IVOA documents page

IVOA Proposed Recommendation

International Virtual Observatory Alliance

IVOA Documents

Simulation Data Access Layer Version 1.0

IVOA Proposed Recommendation 30 January 2017

Interest/Working Group:

<http://www.ivoa.net/wiki/bin/view/IVOA/IvoaDAL>

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Abstract

The Simulation Data Access Layer protocol (SimDAL) defines a set of resources and associated actions to discover and retrieve simulations and numerical models in the Virtual Observatory. SimDAL and the Simulation Data Model are dedicated to cover the needs for the publication and retrieval of any kind of simulations: N-body or MHD simulations, numerical models of astrophysical objects and processes, theoretical synthetic spectra, etc... SimDAL is divided in three parts. First, SimDAL Repositories store the descriptions of theoretical projects and numerical codes. They can be used by clients to discover theoretical services associated with projects of interest. Second, SimDAL Search services are dedicated to the discovery of precise datasets. Finally, SimDAL Data Access services are dedicated to retrieve the original simulation output data, as plain raw data or formatted datasets cut-outs. To manage any kind of data, eventually large or at high-dimensionality, the SimDAL standard lets publisher choose any underlying implementation technology.

Status of this document

This is an IVOA Proposed Recommendation made available for public review.

It is appropriate to reference this document only as a recommended standard that is under review and which may be changed before it is accepted as a full recommendation.



VO-Theory standards

Remise à plat du service d'accès aux vocabulaires pour SimDM / SimDAL

Nicolas Moreau

Zakaria Meliani

David Languignon

Franck Le Petit

Home Search concepts Credits

Help

This service is dedicated to scientists and VO developers who wish to publish theoretical services described by [the Simulation Data Model](#).



As described in the [IVOA](#) standard, Simulation Data Model, registrations of theoretical services, require to provide several URIs corresponding to semantics keywords describing services and simulations. VO-Theory concepts are based on SKOS description as recommended by [the IVOA Semantic Working Group](#).

Example of a VO-Theory URIs : <http://purl.obspm.fr/vocab/Algorithms/GaussSeidel>

This website is dedicated to the discovery of these URIs. Navigate through the broader, narrower, related terms to discover the most precise concept you wish.

To suggest new concepts or corrections, contact : support.votheory@obspm.fr.

Search concepts

IVOA Theory Vocabularies | Specific vocabularies

Physical processes

Concepts

Quick search

AGN Feedback Absorption Acceleration Of Particles Accretion

Advection Alpha Process Astrochemistry Atomic Cooling

Atomic Processes Barotropic Equation Of State

Birkeland Current Sheet Bremsstrahlung Chaos

Chemical Reaction Chemistry Collision Collisional Broadening

Collisional Excitation Collisional Plasma Collisional Processes

Compton Effect Compton Scattering Conduction Convection

Cooling Processes Coronal Mass Ejection Current Sheet

Cyclotron Decaying Turbulence Diffusion Drag Force

Driven Turbulence Dynamics Electrodynamics

Electromagnetism Energy Transfer Equation of State

Fluorescence Forcing Compressive Turbulence

Forcing Solenoidal Turbulence Galactic Wind Feedback

Gas-Grains Collisions General Relativity Geomagnetic Storm

Gravitation Gross Tail Current Sheet Heating Processes

Hydrodynamics Instability Inverse Compton

Inverse Compton Scattering Isothermal Equation Of State

Line Cooling

<http://purl.obspm.fr/vocab/PhysicalProcesses/LineCooling>

Broader concepts

[Cooling Processes](#)

Broader Transitive concepts

[Physical Process](#)

[Cooling Processes](#)

Related concepts

[Atomic Cooling](#)

[Molecular Cooling](#)

Will be discussed at next InterOp

To be used to tag quantities in SimDM