# **VO Science Applications**

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# Turning point into the VO era

- Essential for imminent data volumes and rates
- Multi-  $\lambda$  science requires
  - Data from different telescopes
  - Analysis tools
  - on-line services
  - archived information

#### to be *readily compatible*

VO = framework for interoperable systems

• VO Vision: All Astronomy resources as if they were on your desktop



### Astronomy Data interoperability

- Images (multi-band, mosaics, cutouts, FT)
- Catalogues
- Spectra
- Time series
- Spatial Information
  - Sky regions, slits, FoV, etc.
- Simulation data many kinds
- Multi-d data sets
  - Data cubes, irregularly spaced data
  - Plots, multi-d cluster analysis and correlations

# AVO Science Reference Mission

- Circumstellar disks: from pre-Main Sequence stars to stars harbouring planets
- Intermediate Velocity Clouds
- Which Star will go Supernova next?
- Initial Mass Function within 1kpc: Planetary to Stellar Masses
- Initial Mass Function for Massive Stars
- Contributions of Low and Intermediate Mass Stars to the ISM
- Galaxy Formation and Evolution from z=10 to 0.1
- Build-up of Supermassive Black Holes
- Formation and Evolution of Galaxy Clusters
- Correlation of CMB, radio/mm and optical/NIR Galaxy Surveys

### Galaxy Formation and Evolution from z=10 to 0.1

- When did the 1<sup>st</sup> objects form?
- What are the progenitors of present day massive ellipticals?
- How many massive galaxies at z>1,2,4?
- How do SF and galaxy stellar mass densities evolve?
- Required data
- Deep Multi-wave surveys (GOODS, COSMOS)
- HST+ACS bviz imaging
- SLOAN
- Optical spectroscopy
- MERLIN, GMRT, VLA, ATCA radio
- Chandra and XMM-Newton X-ray
- Spitzer mid-IR
- Future sub-mm
- GALEX UV imaging

"Only now, and only with through the VO, are the datasets large enough, and the tools mature enough that Galaxy Formation and Evolution can be examined in a meaningful way."

### VO steps

- Extract sample from data
  - Perform Sextractor type photometry
    - Cross correlate with images, catalogues, spectra. Crucial that output results are scientifically useable and reliable Matching of PSF, consisitent photometric apertures, treatment of noise, and upper limits
      - Consistency checks like stellar colours
        - Output multi-band catalogue, and colour-colour diagrams
      - Visualize output colour-colour space
    - Photometric z from SEDs (Template SED libraries, extinction curves etc.)
    - → Physical Parameters L, E(B-V), SFR, M/L  $\in$  stellar mass
  - Comparison with star formation scenarios and synthetic spectra
  - Morphological analysis
    - Stack images at same wavelength, or spectra at different redshifts
      - Build average spectra for specific object classes
        - Angular clustering analysis
  - Comparison with mock catalogues from theoretical simulations

### Correlation of CMB, radio/mm and optical//NIR Galaxy Surveys

#### Integrated Sachs-Wolfe Effect

 CMB fluctuations from passage through time varying gravitataional potential

#### Sunyaev-Zel'dovich Effect

- Inverse compton scattering of photons by plasma in the hot intra-cluster medium
- Required data
- WMAP
- Planck
- radio/IR surveys
- X-ray/optical cluster data

Study of full-sky maps from federated archives to disentagle various cosmological and astrophysical effects



### Intermediate Velocity Clouds

- What is the origin of neutral gas clouds moving with unexpected velocities in the galaxy ?
- Required data
- All sky far-IR surveys DIRBE
- IAR, HIPASS 21cm
- SHASSA H



#### VO Steps

Identify regions of excess HI or H To detect IR IVC, remove foreground using HI Check contamination by warm ISM Classify dust rich/poor IVCs Analyse spatial distribution

## We need (1)

- Visual browsing of data and distributed information
- Visualizing heterogeneous data
  - Combining Multi-wavelength data taking into account different:
    - → Units
    - coverage
    - Resolutions/PSF, observing technique

# We need (2)

- Multi-wave cutouts of individual sources
- Generate and visualize SEDs from image, and spectral and catalogue data
  - Taking into account different
    - Beams/apertures (extended sources)
    - → Backgrounds
    - Photometric systems
- Time axis:
  - → Light-curves
  - Multi-epoch imaging

## We need (3)

Compare observations with models
 Virtual observations of models
 Projection of models to observed parameter space
 Spectral fitting/classification

- Colour-colour visualization Tool
- Astronomy functionality alongside visualization
  - Reproject data, correct for extinction, calulate luminosities etc.
  - Visualization requirements ∈ Analysis requirements

# **VO** Tools

- Prototypes
  - VO-enabled existing tools & new tools
- Services
  - VO-enabled with 'translation layer'
  - New services
- Applications Infrastructure
  - Make tools accessible
  - Build your own customized system from VO components



#### AVO prototype based on CDS Aladin

### Catalogue



- X-match
- Visualization
- Direct links to:







### Image Cutout Tool



#### AVO prototype based on CDS Aladin



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### Simple Spectrum Access

Image / Spectrum / Catalog interoperability

STScI Specview & AVO prototype

# Multi-archive spectra

- SSA servers
- Registry
- Unit interop.



ESA VOSpec



#### Line lists integrated into the tool

**STScl Specview** 

#### Line lists from a Service

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ESA VOSpec & Spectral Line Access Protocol (SLAP) Service

### VisIVO : Visualization tools



INAF, CINEA : visivo.cineca.it





Use Mouse to Rotate Graph, Use Keyboard arrow keys for Zoom-In and Zoom-Out

8666 Points Developed By VO-India.



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### AstroGrid : scaling up VO processing into WorkFlows













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### **IVOA Applications Interest Group**

- Announcements of new tools
- Discussion on VO tools
- Suggestions for enhancements
- Feedback to IVOA working groups on standards etc.

### Working list of applications - add your own!

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## Summary

 VO tools are ready for doing science Varying levels of stability and maturity VO services are rapidly coming online Data centres implementing translation layers Specialized services uptake of VO standards Workflows and Grid capabilities First steps are promising

