

Science in the Virtual Observatory

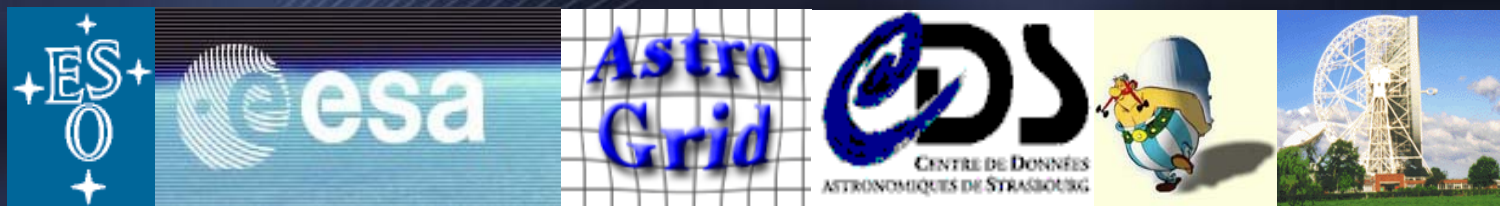
Paolo Padovani

*Head, ESO VO Systems Department
EURO-VO Facility Centre Scientist*

- VO as a science driver
- VO-Science in Europe: the AVO experience
- The AVO Science Reference Mission
- The future: science in the EURO-VO

- An innovative, *evolving* system, which takes advantage of astronomical data explosion
- It will allow users to interrogate multiple data centres in a seamless and transparent way and to utilize at best astronomical data
- It will allow new *SCIENCE* by, e.g.,:
 - ✓ moving Astronomy beyond era of “classical” identification by combining **all** available information: data mining (increase observational efficiency) + statistical identification (less need for spectra)
 - ✓ permitting massive, remote computing on astronomical data, with analysis tools residing where the data are or in a VOStore

- Astrophysical Virtual Observatory Project: R&D on scientific requirements and technology for building the VO in Europe, 50% funded by European Community



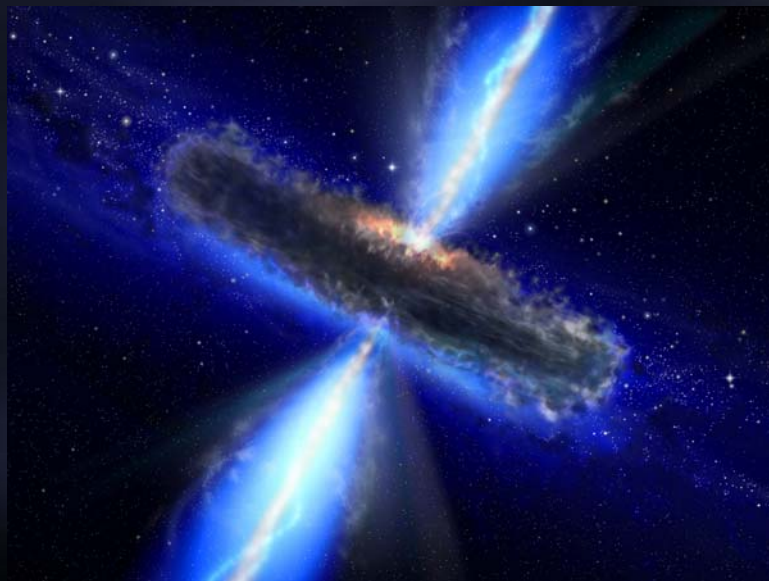
- Phase A, 2001 - 2004/5 (<http://www.euro-vo.org>)
- Driven by strategy of **scientific VO annual demonstrations**
- Science Working Group established to provide scientific advice to AVO project
- Project completed; now EURO-VO

- Two scenarios:
 - Extragalactic: Obscured (Type 2) Quasars
 - Galactic: Classification of Young Stellar Objects (YSO)
- Multiwavelength, heterogeneous, and complex data: VLA, CGPS, ISO, 2MASS, USNO, 2.2m/WFI, VLT/FORS, HST/ACS, XMM, and Chandra (images, spectra, and catalogues)
- Access to any **VO-compliant** data: seamless and transparent access to ESA ISO & XMM archives and ESO data products
- **AVO: from First Light (2003 demo) to First Science!**

VO First Science!

- First refereed astronomical paper enabled via end-to-end use of VO tools and systems:
"Discovery of optically faint obscured quasars with Virtual Observatory tools",
Padovani, Allen, Rosati, & Walton, 2004, A&A, 424, 545

ESA/ESO Press release May 28 2004

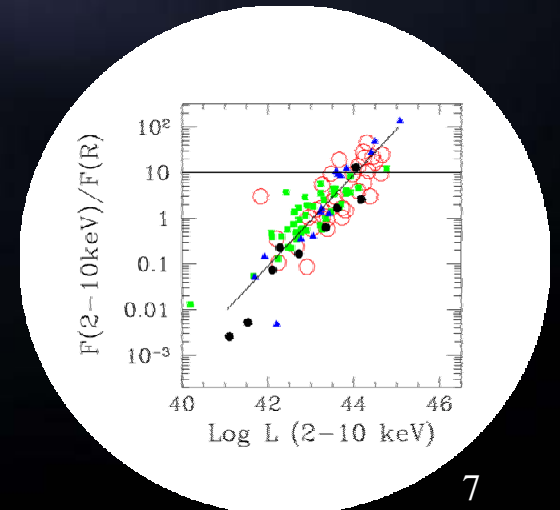
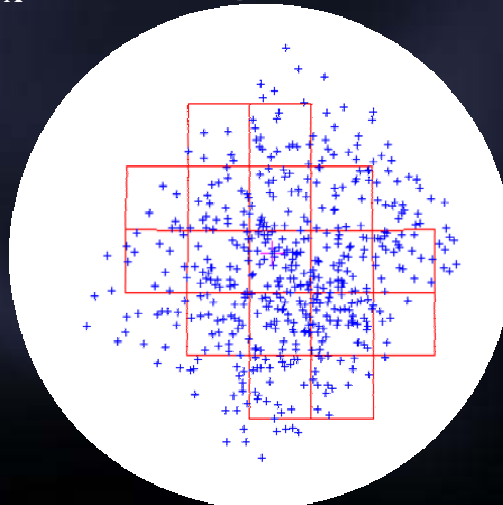


- Evolution of Aladin (Centre de Donnés astronomiques de Strasbourg [CDS])
- Downloadable Java application
(<http://www.euro-vo.org/twiki/bin/view/Avo/SwgDownload>)
- Extensible toolset with plug-ins which allows easy access to images (manipulation), spectra, catalogues, with overlays, plotting facilities, catalogue extraction, and a cross-correlation utility [ASTROGRID, VOIndia, STScI]
- Interoperable with other VO tools

Discovery of QSO 2s with VO tools

- GOODS HST/ACS data & catalogues
- Chandra X-ray catalogues
- Select absorbed X-ray sources
- Cross-match X-ray and optical
- Check against spectroscopy
- Apply empirical estimator for L_x :

$L_x > 10^{42}$ erg/s: AGN 2; $L_x > 10^{44}$ erg/s: QSO 2



2MASS
ESO-WFI
Chandra
VLT-ISAAC
HST-ACS
DSS
My Data

Data Tree

- GOODS-WFI
 - V89
 - DEEPC-FV-Preview 38.1 'x37.3' 2000-10-2
 - DEEPC-FV 8.2 'x8.2' 2000-10-26
 - B99
 - RC162
 - U38
- GOODS-ACIS
 - LR.1-10KEV
 - ACISHCDFSN000 1.2 'x1.2' 1999-10-14
 - HR.1-10KEV
- GOODS-ISAAC
 - J
 - GOODS-10 2.5 'x2.5' 08/04/2002
 - GOODS-11 2.5 'x2.5' 08/04/2002
 - GOODS-14 2.5 'x2.5' 08/04/2002
 - GOODS-15 2.5 'x2.5' 08/04/2002
 - GOODS-20 2.5 'x2.5' 08/04/2002
 - GOODS-16 2.5 'x2.5' 08/04/2002
 - GOODS-21 2.5 'x2.5' 08/04/2002
 - GOODS-9 2.5 'x2.5' 08/04/2002
 - H
 - KS
- GOODS-HST-ACS
 - F775W
 - epoch1
 - epoch2
 - epoch3
 - epoch4
 - epoch5
 - version1.0
 - CDF-SOUTH-SECT32-VERSION1.0
 - CDF-SOUTH-SECT25-VERSION1.0
 - CDF-SOUTH-SECT23-VERSION1.0
 - CDF-SOUTH-SECT21-VERSION1.0
 - CDF-SOUTH-SECT44-VERSION1.0
 - CDF-SOUTH-SECT14-VERSION1.0
 - CDF-SOUTH-SECT42-VERSION1.0
 - CDF-SOUTH-SECT12-VERSION1.0
 - CDF-SOUTH-SECT35-VERSION1.0
 - CDF-SOUTH-SECT33-VERSION1.0
 - CDF-SOUTH-SECT31-VERSION1.0
 - CDF-SOUTH-SECT24-VERSION1.0
 - CDF-SOUTH-SECT22-VERSION1.0
 - CDF-SOUTH-SECT45-VERSION1.0
 - CDF-SOUTH-SECT43-VERSION1.0
 - CDF-SOUTH-SECT13-VERSION1.0
 - CDF-SOUTH-SECT11-VERSION1.0
 - CDF-SOUTH-SECT34-VERSION1.0
 - F606W
 - F435W
 - F850LP
- SERC
 - J
- AAO

Data available at selected point are highlighted in tree

Info Frame

CDF-SOUTH-SECT23-VERSION1.0

Observation_Name	CDF-SOUTH-SECT23-VERSION1.0
ObservingProgram_Name	GOODS-HST-ACS
FilterName	F775W
Size_alpha	4.1'
Size_delta	4.1'
Angular Pixel Size	0.029"
Origin	STSCI
OriginalCoding	FITS
CentralPoint_RA	03:32:38.72
CentralPoint_DEC	-27:48:18.3
DateAndTime	2002-08-01
Position Angle	0.0°

Cutout Target: 03 32 33.50 -27 47 36. Grab

Stick FoV in stack LOAD Close

Image metadata

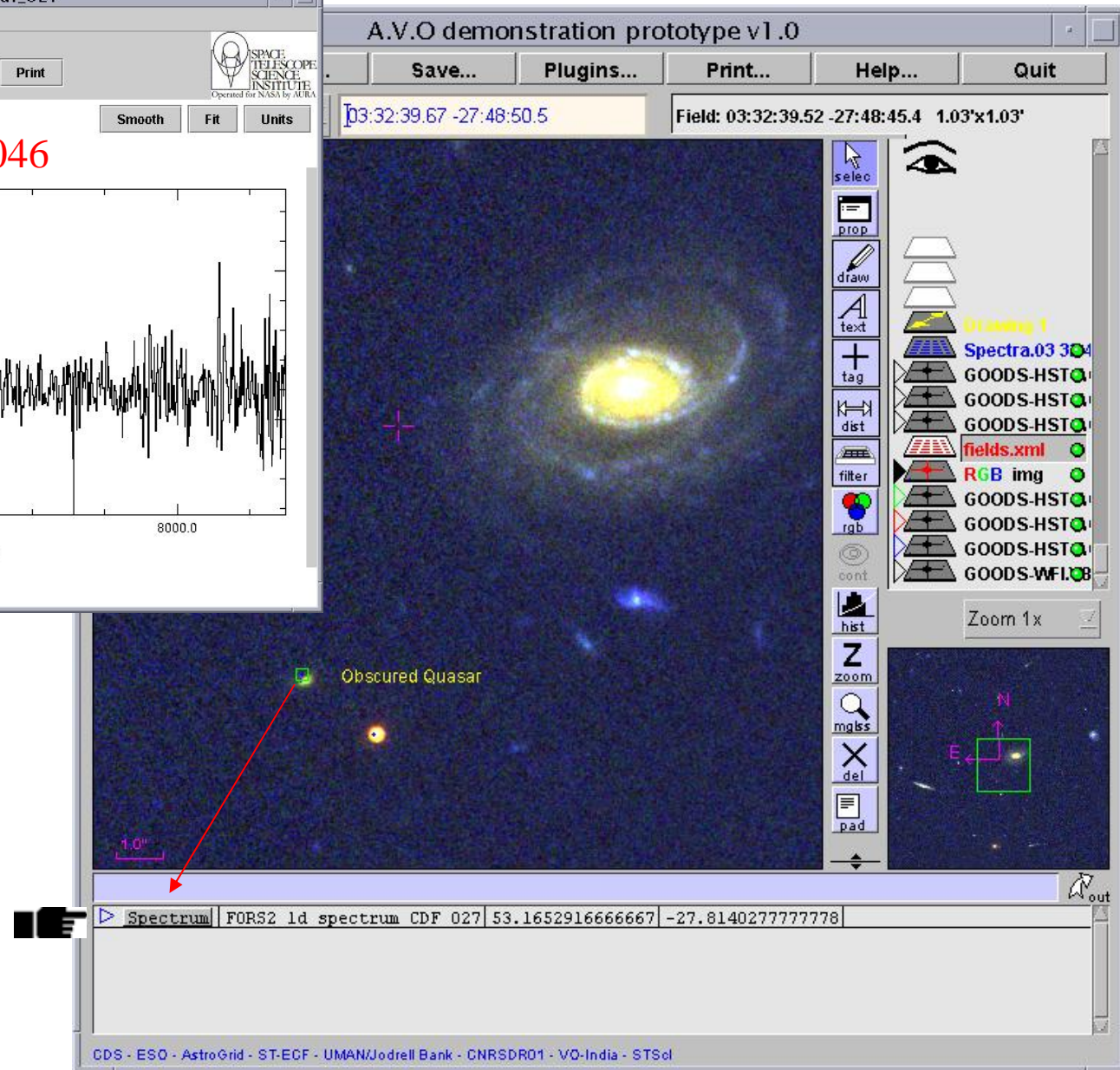
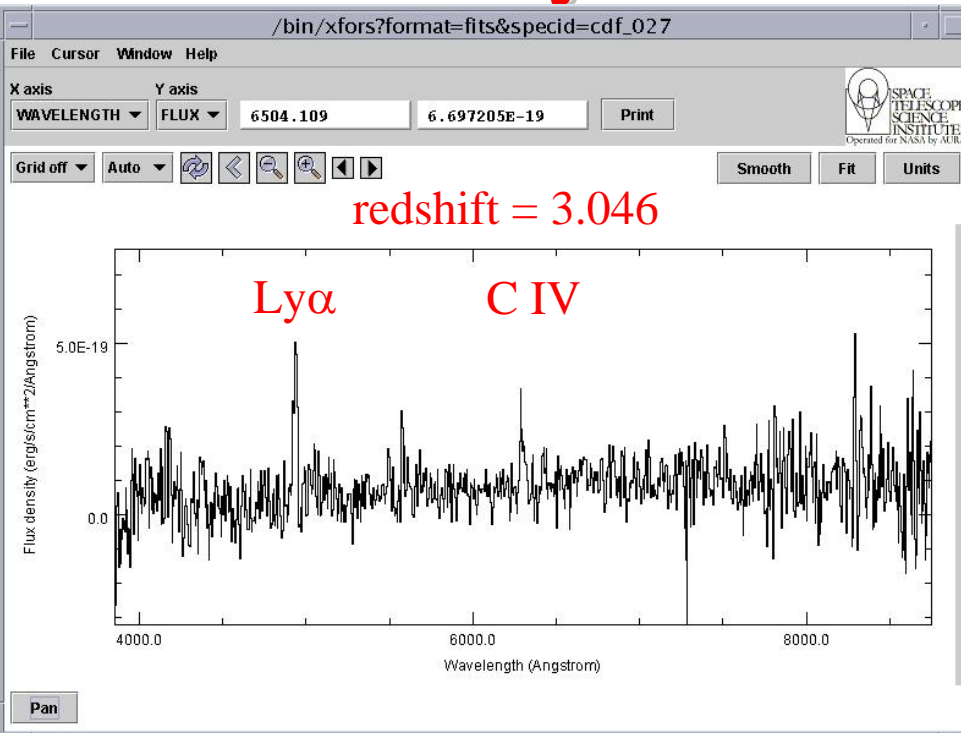
A.V.O demonstration prototype v1.0

Field: 03:32:25.77 -27:48:07.4 38.08'x37.2'

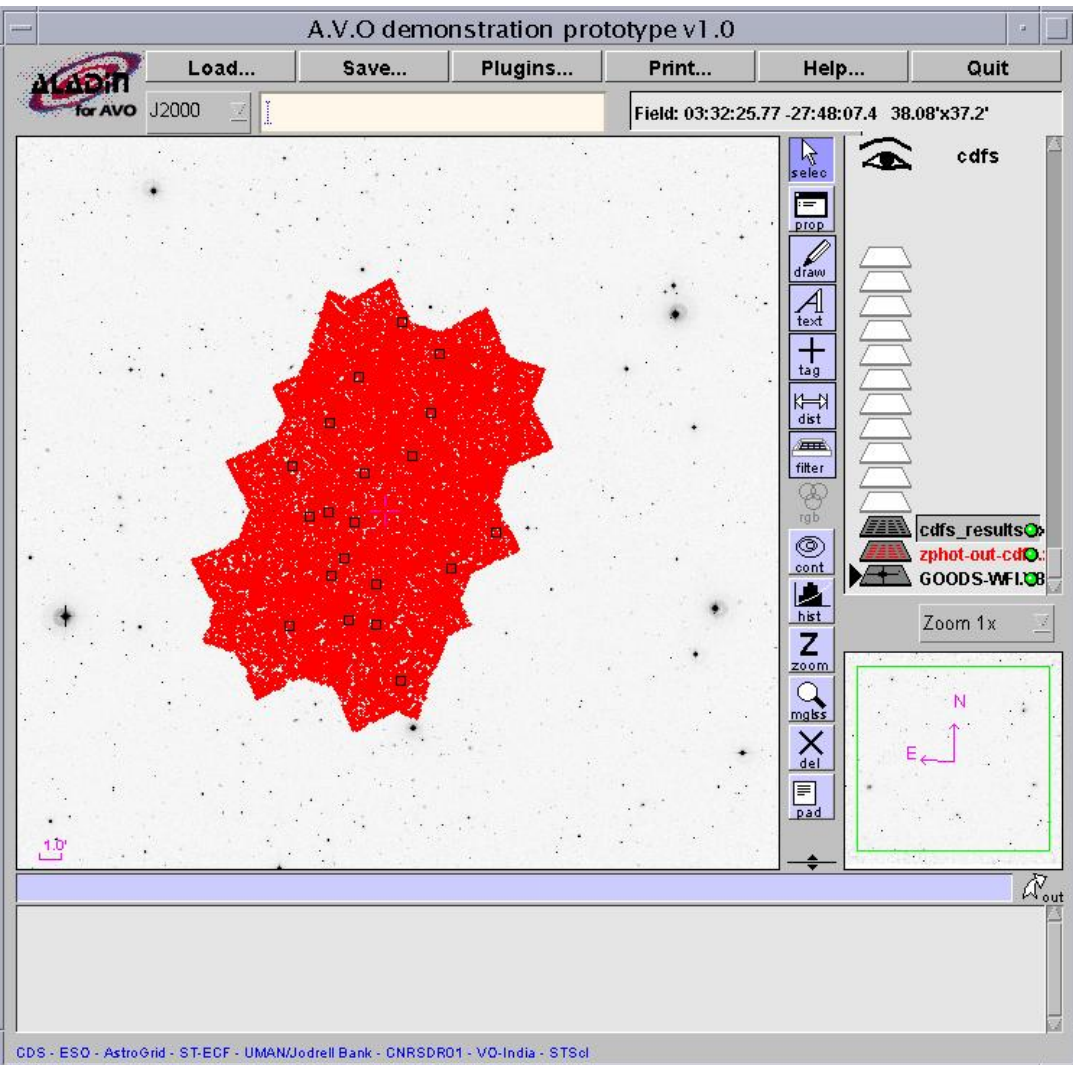
Field of view outlines are plotted automatically

Field of view outlines are plotted automatically

January 2004: AVO First Science



January 2004: AVO First Science



31 new QSO 2 (68
new type 2 AGN) in
GOODS CDFS+HDFN

x 4 increase!

$N(f_x > 10^{-15} \text{ c.g.s.}) \geq$
 $330/\text{deg}^2$

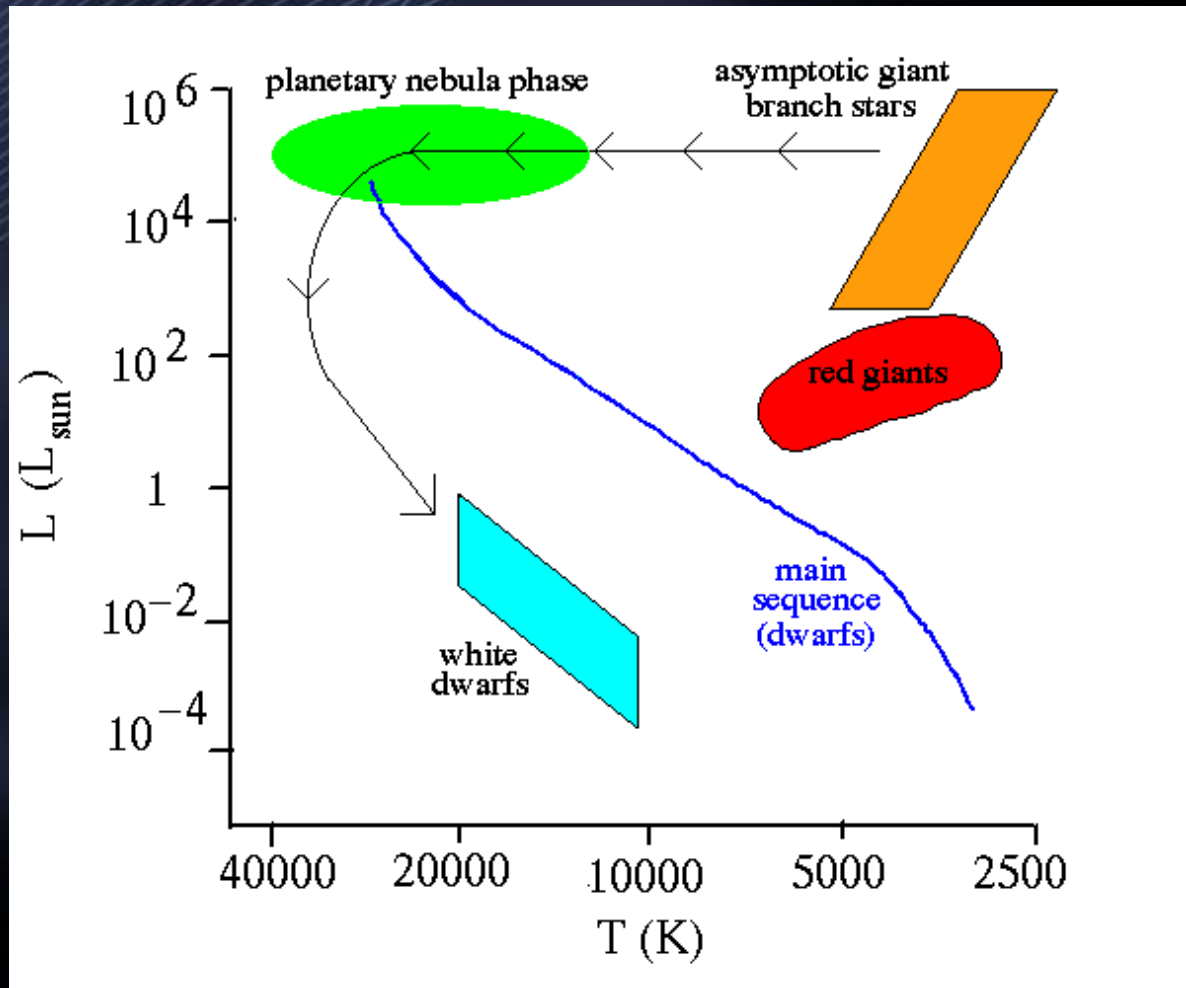
faint: $\langle R \rangle \sim 27$,
beyond spectroscopy
even at VLT/Keck

- AVO has done cutting-edge science by exploiting the data beyond “classical” identification limits ($R > 25$)
- AVO provides “statistical” identification of sources using multiwavelength information
- VO tools enable astronomers to reach into new areas of parameter space with little effort
- “AVO should enable everyone to compete with the GOODS team (on their data)”
[G. Gilmore, SWG meeting, June 2002]

- Two scenarios:
 - Galactic: Asymptotic Giant Branch to Planetary Nebulae Transition (strong science case, Spectral Energy Distribution building from archival data)
 - Extragalactic: Star Formation Histories in Galaxies (new VO computing concepts: towards the Grid; access to theoretical models)
- Multiwavelength, heterogeneous, and complex data: VLA, MERLIN, Spitzer, ISO (spectra and images), MSX, IRAS, 2MASS, WFS, HST/WFPC2, FUSE, IUE, plus VizieR catalogues
- Another VO paper to be written (Bayo, Garcia-Lario, Sierra, et al.)!

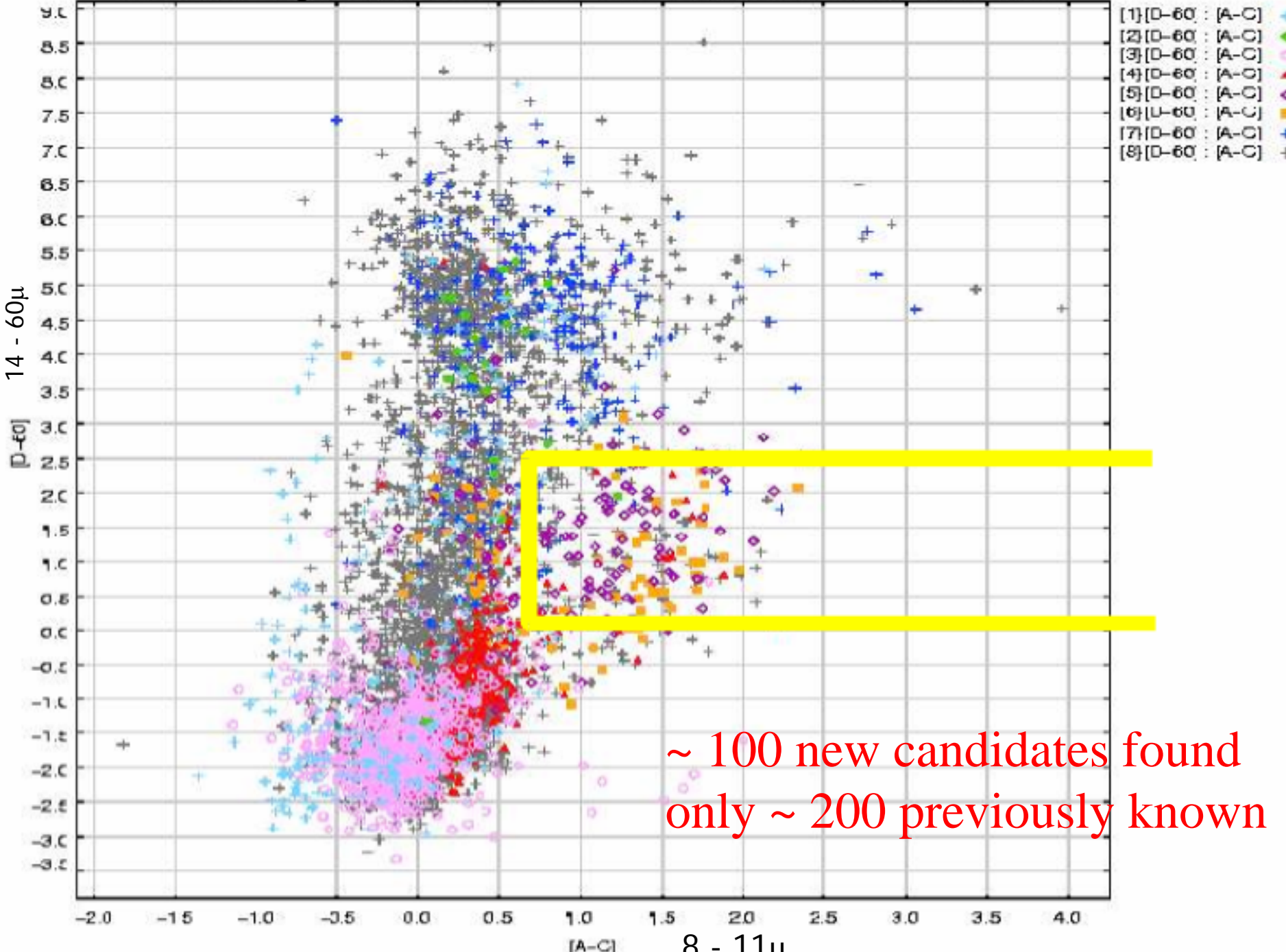
Late evolutionary stages of low- and intermediate mass stars ($1 - 8 M_{\odot}$)

- Short transition times ($\sim 10^3 - 10^4$ yr)
- Recent AGB mass loss (up to $10^{-4} M_{\odot}/\text{yr}$)
- ⇓
- Few objects in this phase ($\approx 1/1,000,000$)
- Many are heavily obscured in the optical
- ⇓
- Need for systematic surveys using infrared data



- Strategy: use large IR catalogues (IRAS and Midcourse Space Experiment [MSX]) to identify colours characterising various classes of sources as given by SIMBAD and then find objects with no published identification in the region of parameter space typical of transition sources.
- Steps:
 - Select high-quality detections in IRAS (~ 70k) and MSX (~ 50k) catalogues
 - Cross-correlate MSX and IRAS lists
 - Calculate 8 - 11 μ and 14 - 60 μ colours
 - Query SIMBAD to get source classification
 - Identify region in two-colour diagram dominated by post-AGB and PNe objects
 - Unclassified sources in the same region \Rightarrow new transition candidates

1=Stars, 2=Extragalactic, 3=AGB, 4=OH/IR, 5=Post-AGB, 6=PNe, 7=YSO, 8=Unclassified



The AVO Science Reference Mission

- The Science Reference Mission (SRM) defines key scientific results that the full-fledged (Phase B) EURO-VO should achieve when fully implemented.
- It consist of ten science cases, with related requirements, against which the success of the EURO-VO will be measured.
- Put together with input from the AVO Science Working Group.

AvoSRM**AVO TWiki system**

Logged in as
[PaoloPadovani](#)

AVO
 SWG
 Management
 Science
 Standards
 Technology
 Reports&Minutes
 Forum
 Events
 Contacts
 Help

EGG
 Main
 Know
 Test
 TWiki

The AVO Science Reference Mission

What is the Science Reference Mission (SRM)? The SRM is a definition of the key scientific results that the full-fledged, Phase B, EuroVO should be able to achieve when fully implemented. It will consist of a number of science cases, with related requirements, against which the success of the EuroVO will be measured.

Contents:

- [SRM document](#)
- [SRM Cases](#)
- [Science Requirements from Euro-VO Partner Projects](#)
- [Background Material](#)

SRM document

- [preliminary version \(.pdf\)](#)

SRM Cases

- [Circumstellar Disks \(.txt\)](#)
- [Intermediate Velocity Clouds \(.txt\)](#)
- [Which Star will go Supernova next? \(.txt\)](#)
- [Initial Mass Function \(low masses\) \(.txt\)](#)
- [Initial Mass Function \(high masses\) \(.txt\)](#)
- [Low and intermediate mass stars contribution to the ISM \(.txt\)](#)
- [Galaxy Formation and Evolution \(.txt\)](#)
- [Build-up of Supermassive Black Holes \(.txt\)](#)
- [Formation and Evolution of Galaxy Clusters \(.doc\)](#)
- [Correlation of CMB background, radio/mm, and optical/NIR Galaxy Surveys \(.txt\)](#)

- EURO-VO funding situation still not well-defined, only VO-Tech (a VO-TC project) funded
- There will be a EURO-VO external science body, whose composition will soon be finalized
- More complex role than AVO SWG given the tri-partite nature of EURO-VO: Data Centre Alliance, Technical Centre, and the Facility Centre

- AVO has reached out to the European astronomical community mostly through the SWG, plus papers and articles (ESO Messenger, ST-ECF & EAS Newsletters)
- EURO-VO will reach out to the data providers through a dedicated workshop at ESO (June 27 - July 1)
- We also plan to start having a presence at JENAM/EAS meetings, to contact directly (and get feedback from) astronomers (as the NVO does at the AAS)

- The Virtual Observatory is a science driver!
- One of AVO's top priorities has been the pursuit of astronomy through:
 - Three science demonstrations
 - Journal and conference astronomical papers
 - A Science Reference Mission
- AVO has produced REAL science tools, which can be (and are!) used for astronomical research:
 - AVO prototype is a very flexible toolset with plug-ins which allow easy access to images, spectra, and catalogues
 - ASTROGRID's workflow builder and MySpace are at cutting edge of VO development
- EURO-VO's plans not fully defined yet; but science will certainly play a major role!